



DLR TAU Results

AIAA DPW-4

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Content

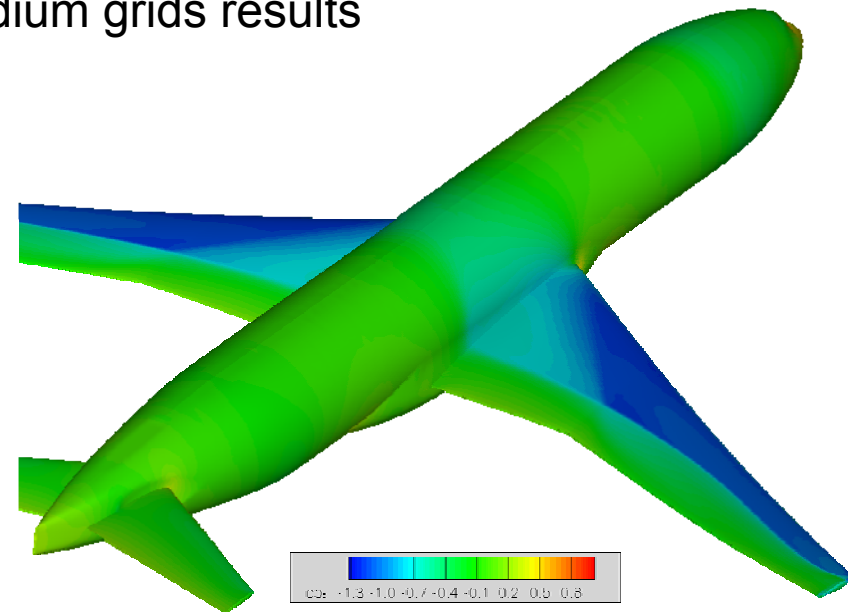
- Objectives
- Grids
- TAU RANS Solver
- Case 1.1: Grid Refinement Study
- Case 1.2: Downwash Study
- Case 3: Reynolds Number Study
- Conclusions



Objectives

— DLR Objectives in DPW-4 —

- Test DLR-TAU with new Solar grid generation approach (hex-dominant in boundary layer):
 - Refinement studies using Solar grids
- Compare to standard TAU Centaur medium grids results (prism-dominant in boundary layer)
- Application of SA, Menter $k\omega$ -SST, and RSM turbulence models
- i_h trim interpolation vs. HTP setting modification in CFD loop (mesh deformation)





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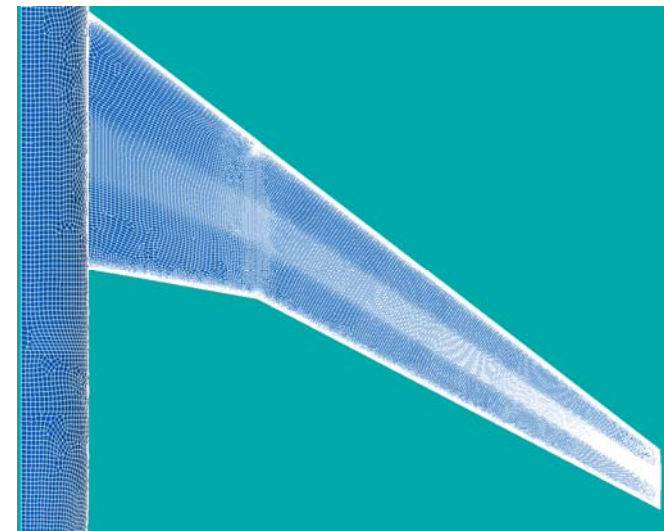
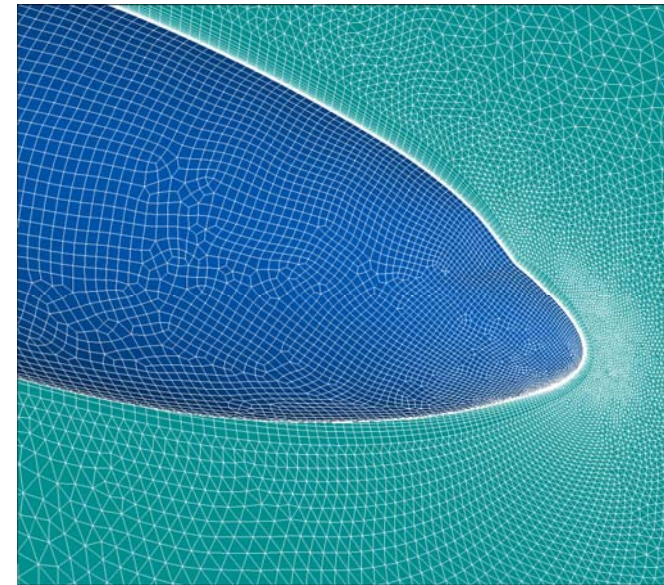
Grids

— Solar —

Solar grid generation system developed by ARA, BAE Systems, Airbus, QinetiQ:

- Anisotropic quad-dominant unstructured surface meshes
- Advancing layer near field mesher
- Buffer layer transitioning to triangulated near field shell
- Tetrahedral far field meshing
- Consistent grid family
- Here: grid refinement factor ≈ 1.42 used

	Coarse	Medium	Fine
Nodes	4074967	11696804	34076798
Hexa Layers	30	42	60



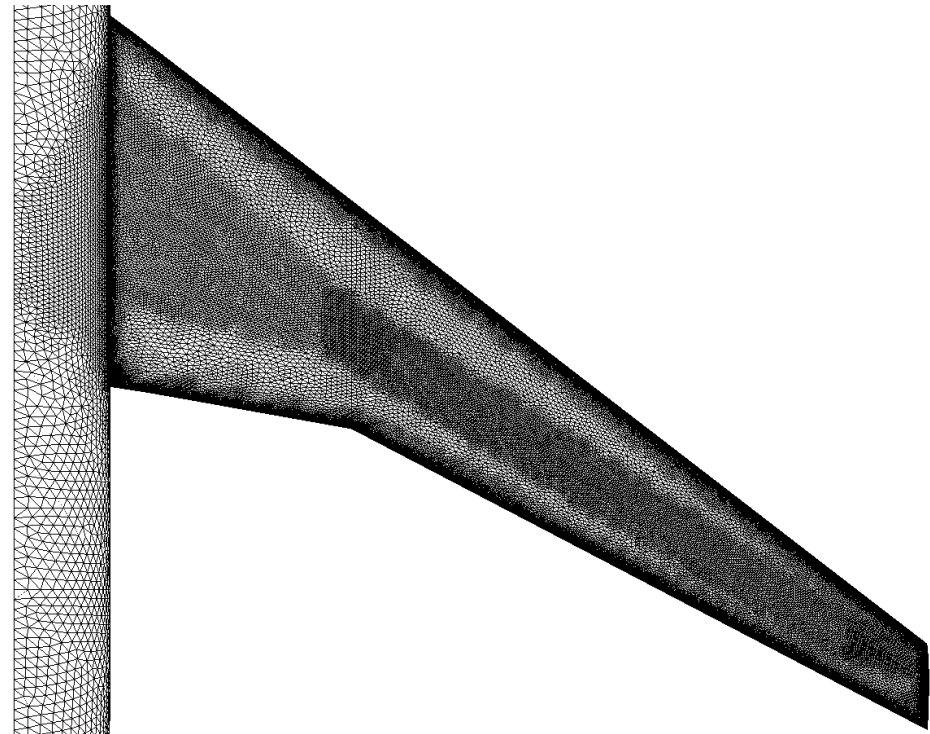
Grids

— Centaur 8.1 —

Centaur grid generation system developed by CentaurSoft:

- Triangulated surface meshes
- Prismatic elements for boundary layer resolution
- Tetrahedral far field meshing

	Medium
Nodes	13331301
Prismatic Layers	35

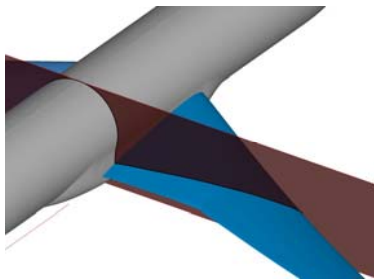
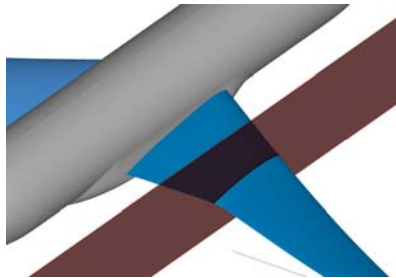




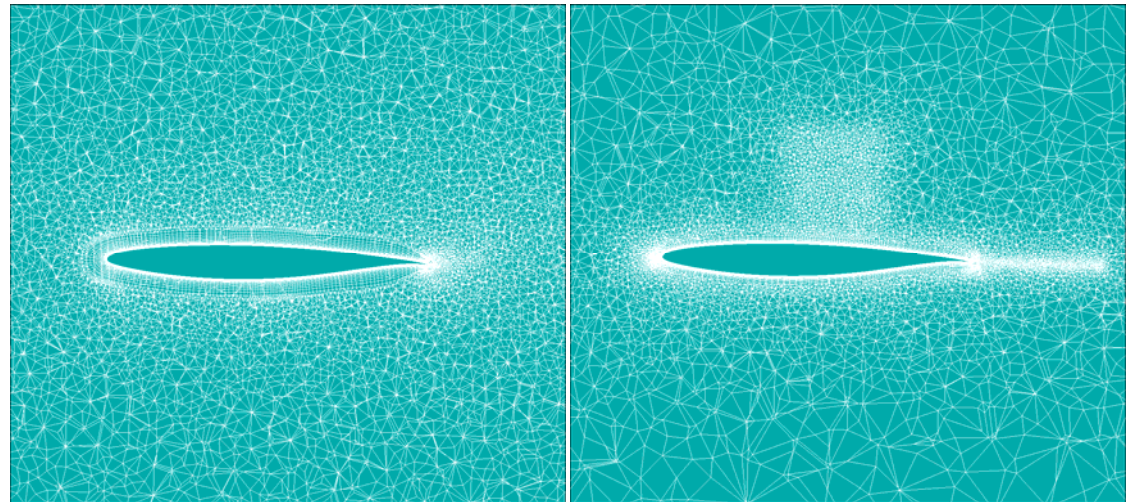
Grids

— Volume Grid —

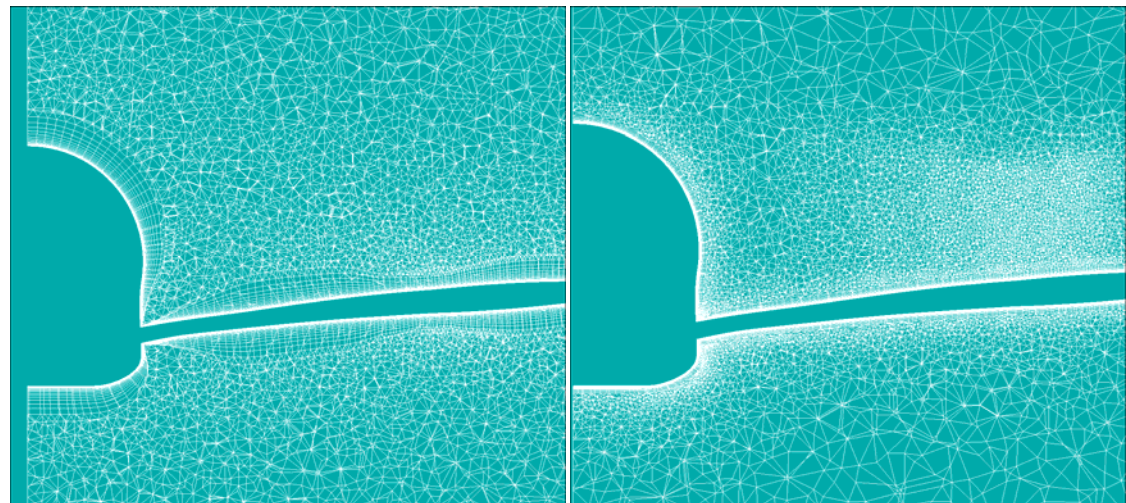
- Medium grids
- Best practice for Centaur grids
- Limited experience with Solar grids



Y=400



X=1400





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TAU RANS Solver

— Overview —

- TAU solves Reynolds-averaged Navier-Stokes equations
- Finite Volume Method, node-centered, (cell-centered), dual grid technique
- Several discretization schemes, here:
 - 2nd order central with Jameson-type dissipation
- Time integration: Runge-Kutta, Backward Euler
- Local time stepping, residual smoothing
- Multigrid
- Several turbulence models, here:
 - Spalart-Allmaras original (SA, SAO)
 - Menter $k\omega$ -SST ($k\omega$ -SST)
 - Speziale-Sakar-Gatski/Launder-Reece-Rodi, SSG/LRR- ω , (RSM)



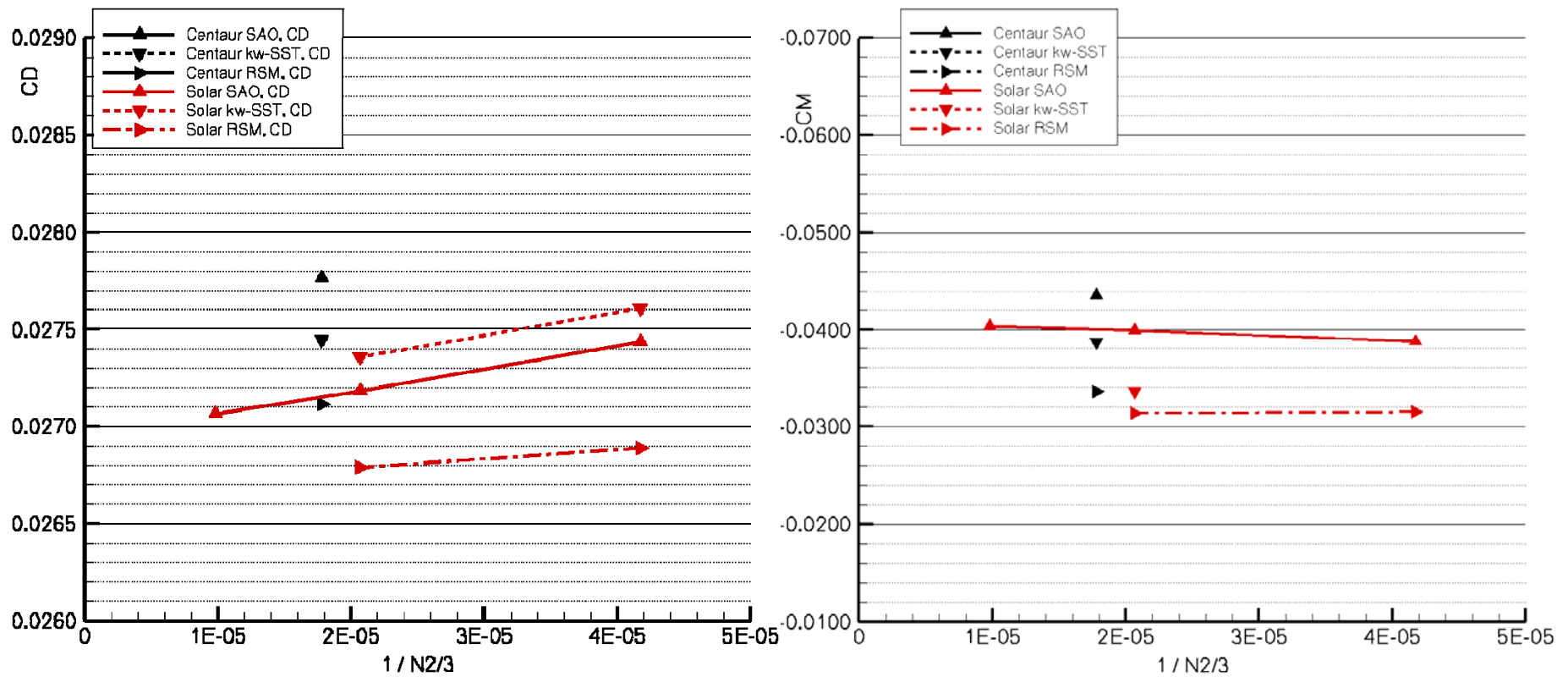
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Case 1.1

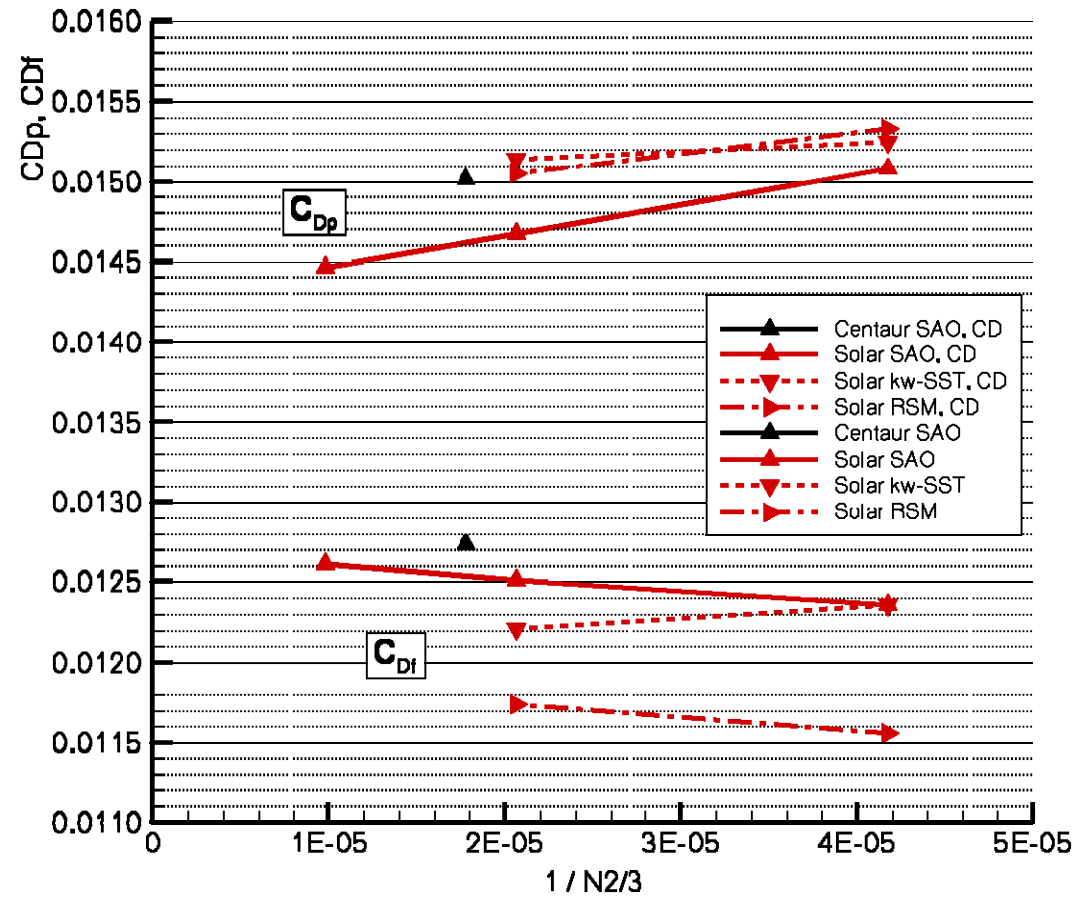
— Grid Type/Size, Turbulence Model —





Case 1.1

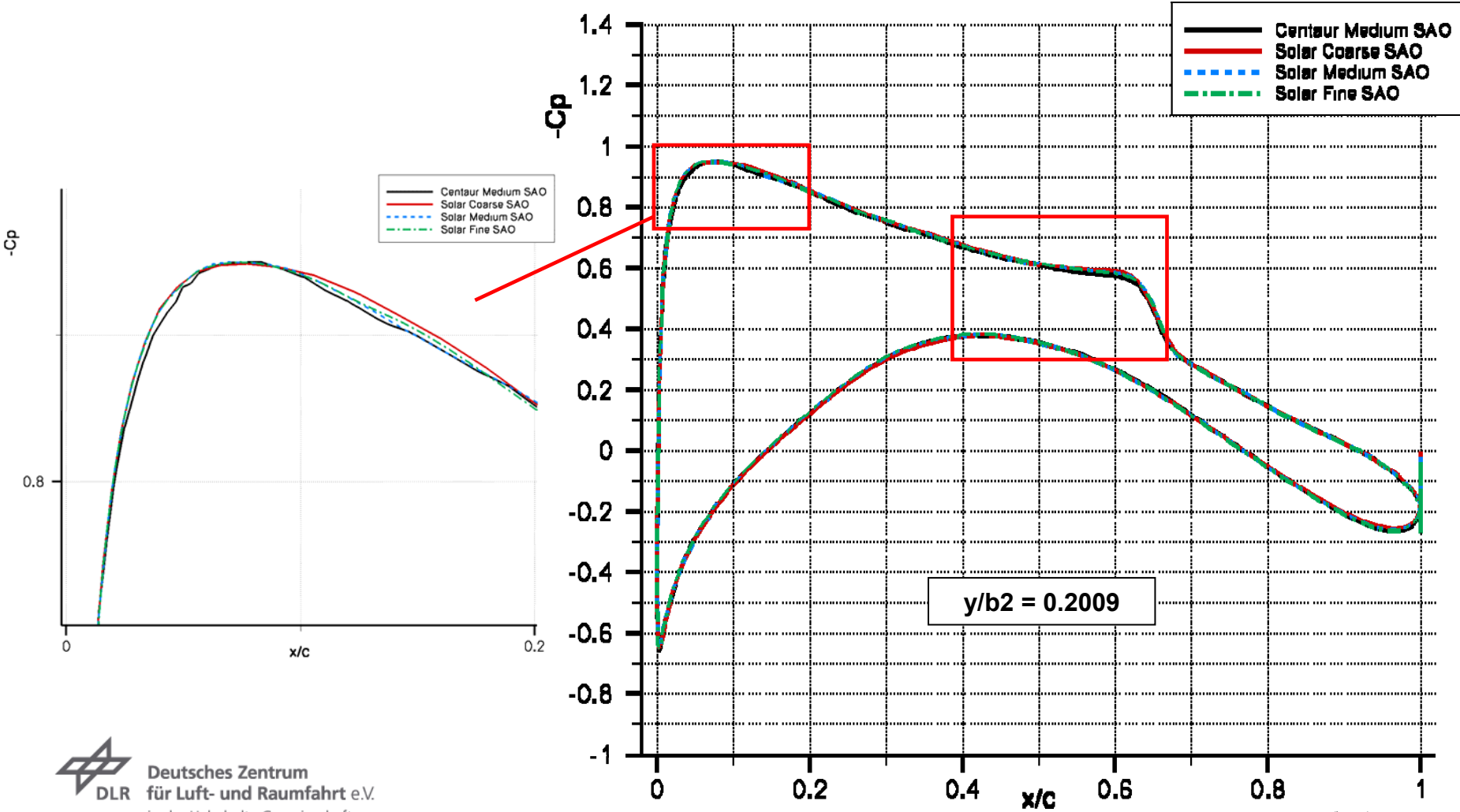
— Grid Type/Size, Turbulence Model —





Case 1.1

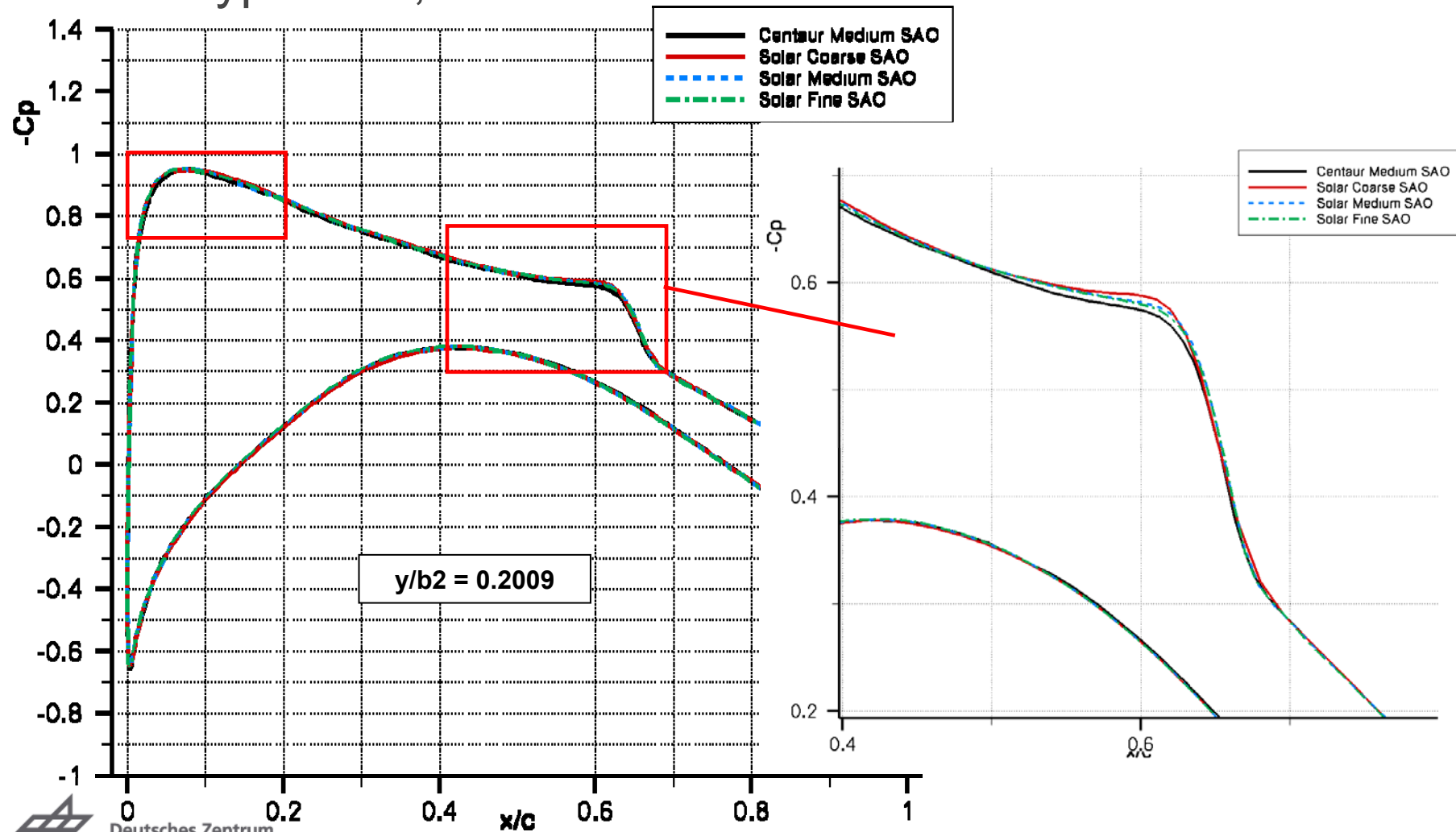
— Grid Type/Size, SAO Model —





Case 1.1

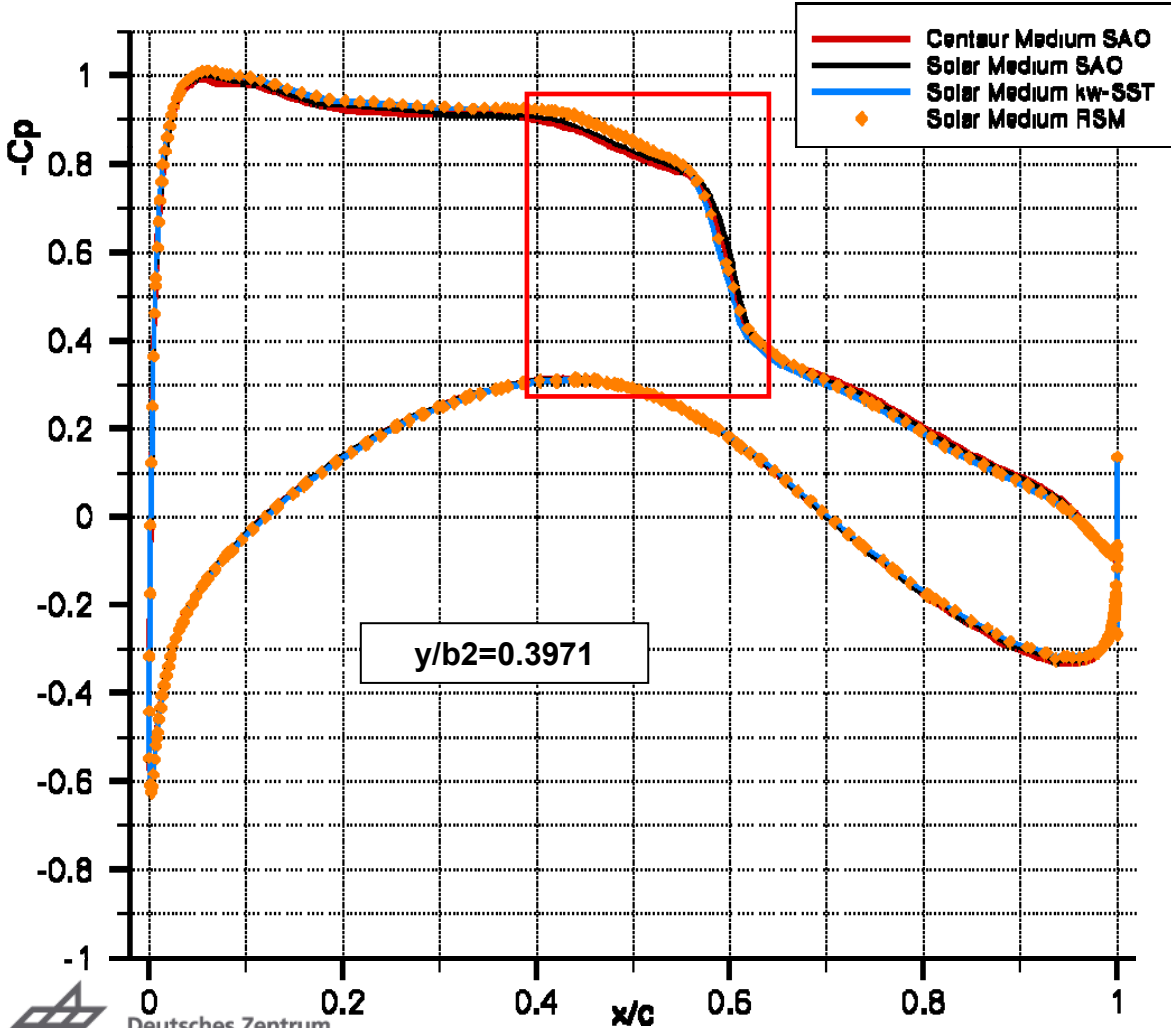
— Grid Type/Size, SAO Model —





Case 1.1

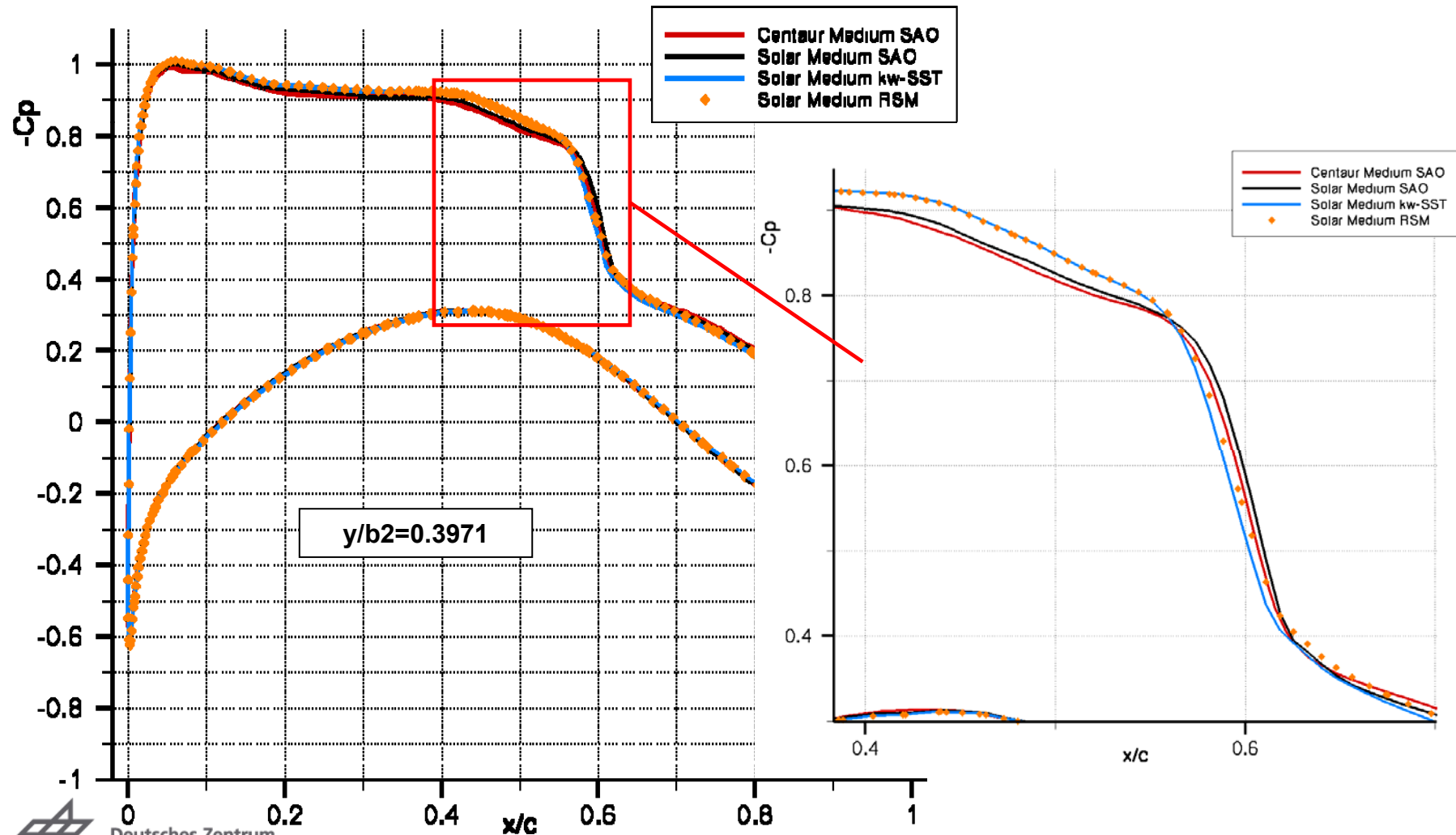
— Turbulence Model —





Case 1.1

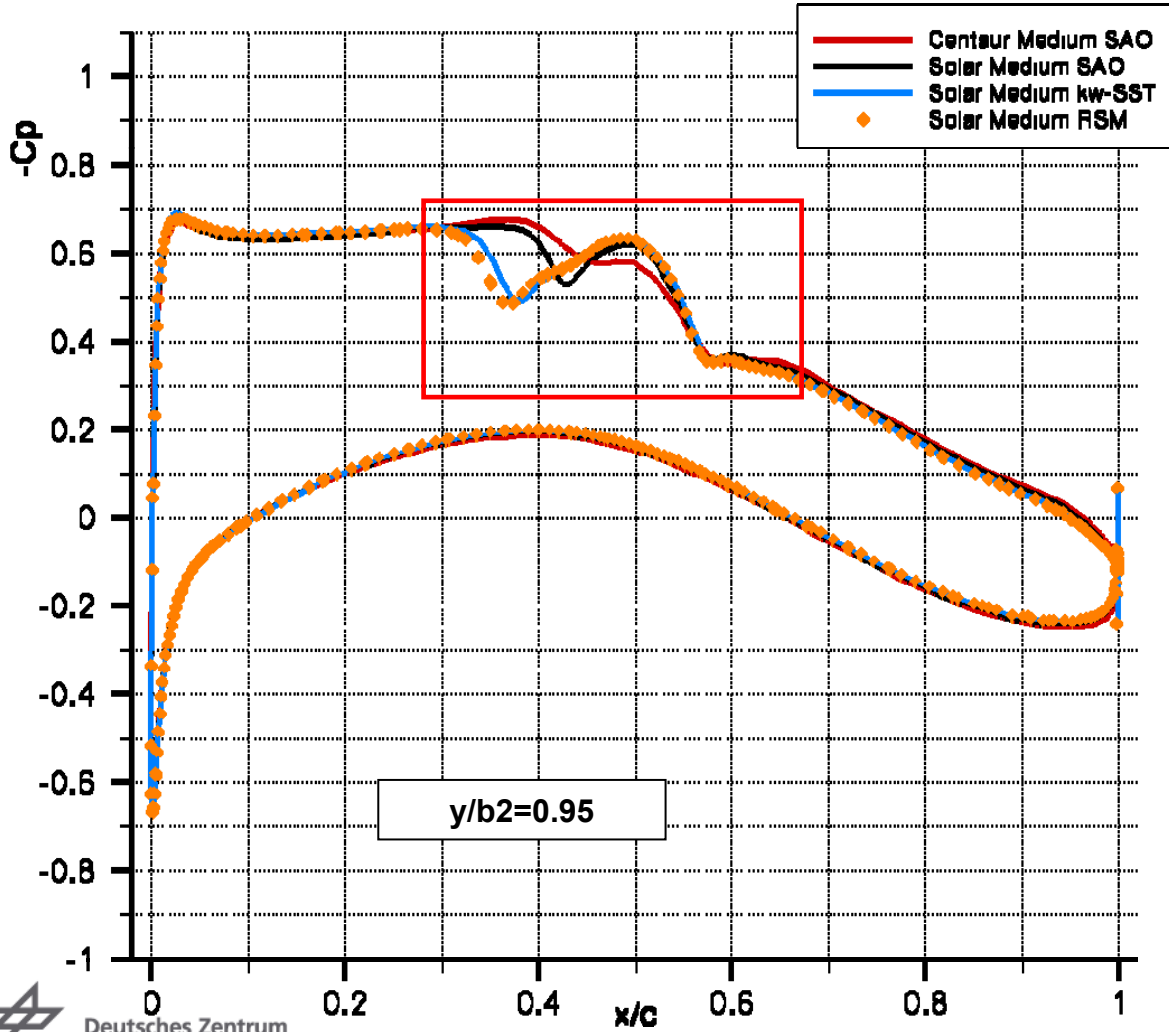
— Turbulence Model —





Case 1.1

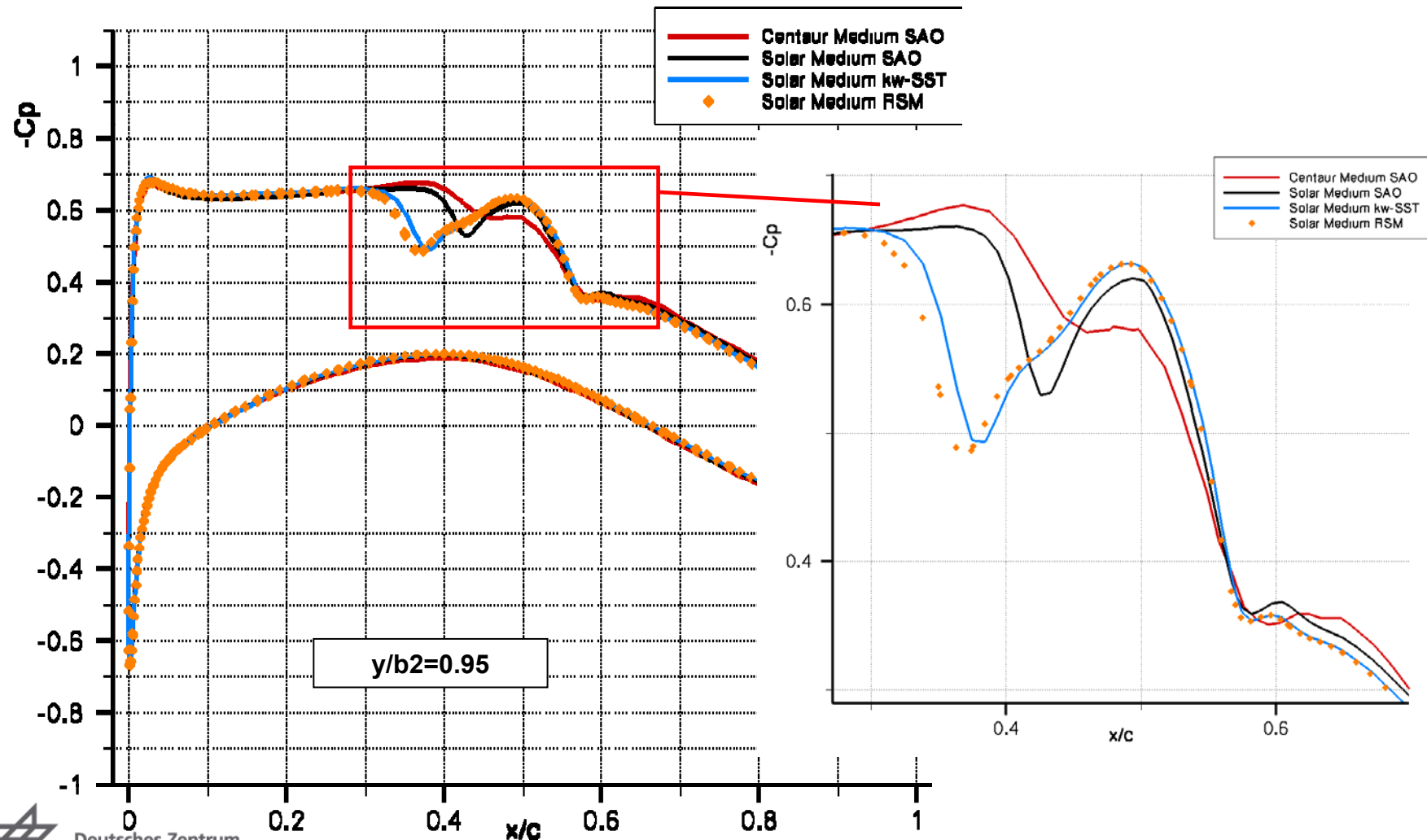
— Turbulence Model —





Case 1.1

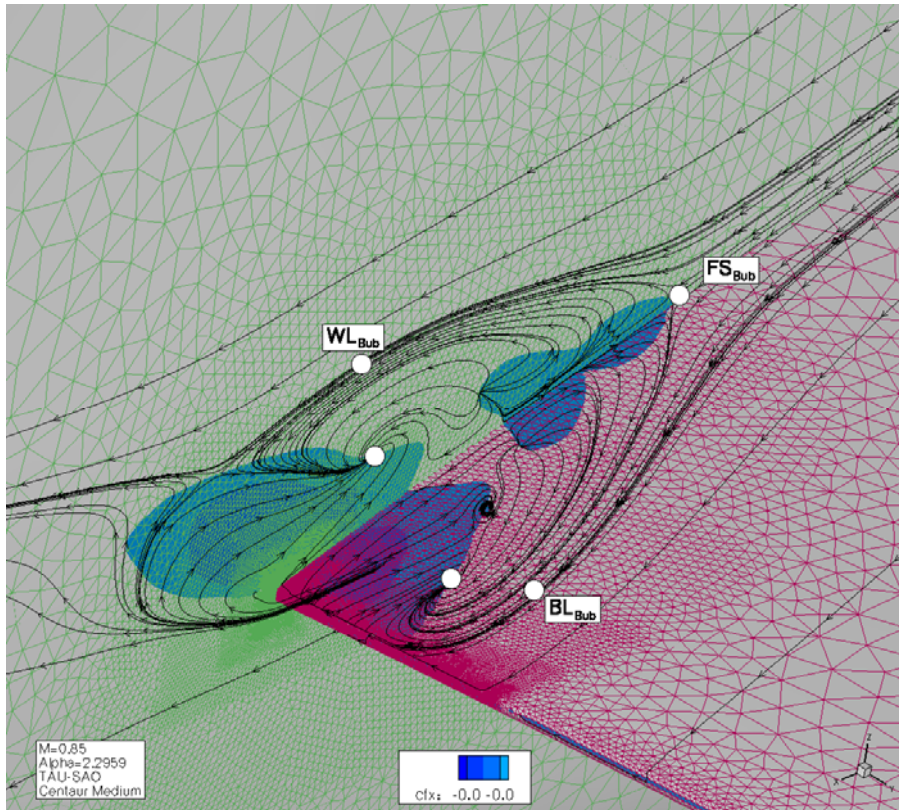
— Turbulence Model —



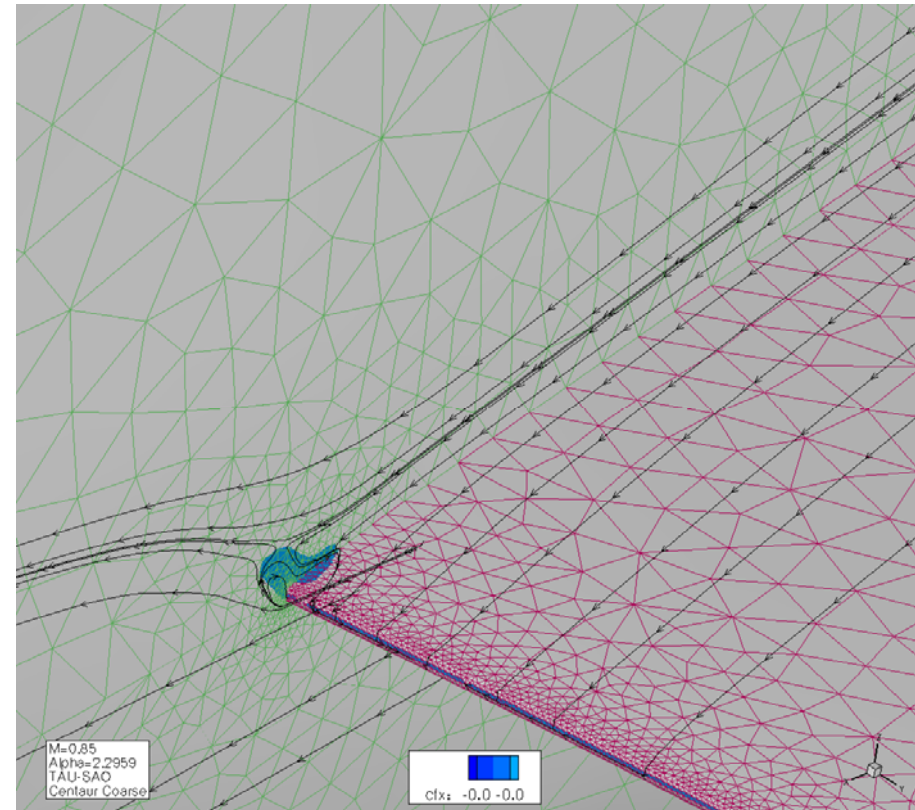


Case 1.1

— Grid Size —



SAO, Centaur grid, medium

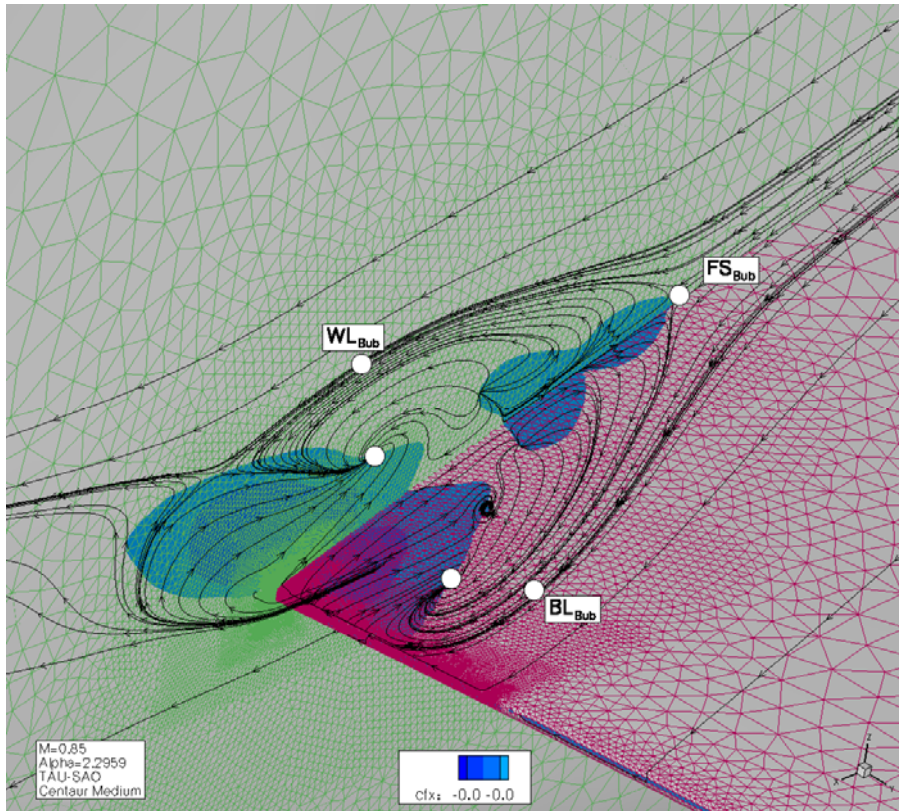


SAO, Centaur grid, coarse

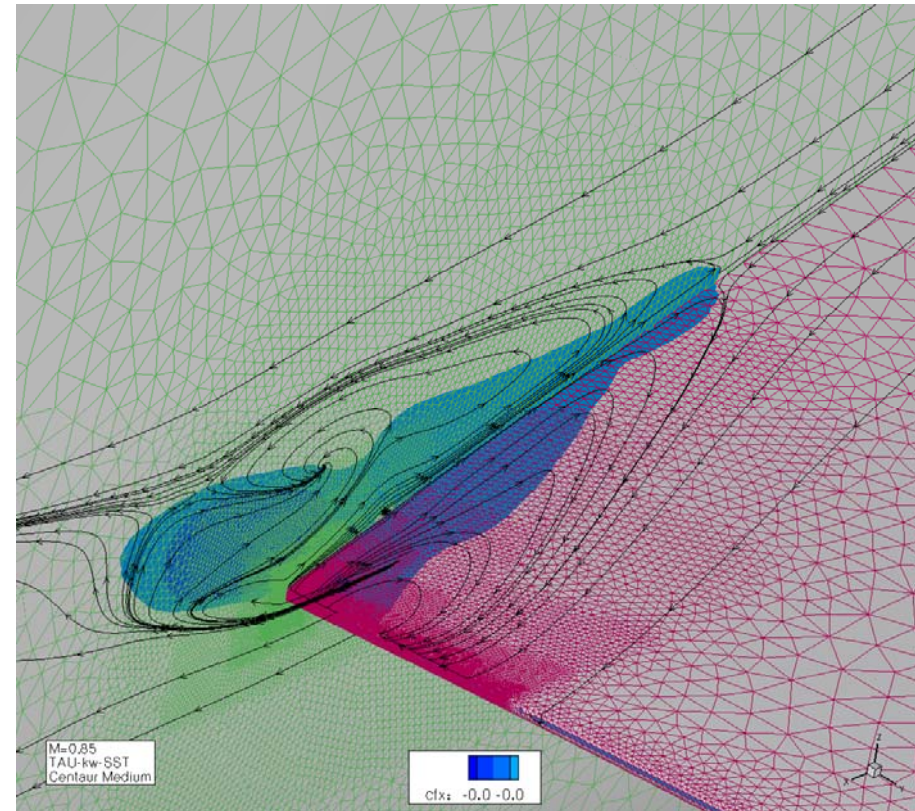


Case 1.1

— Turbulence Model —



SAO, Centaur grid, medium

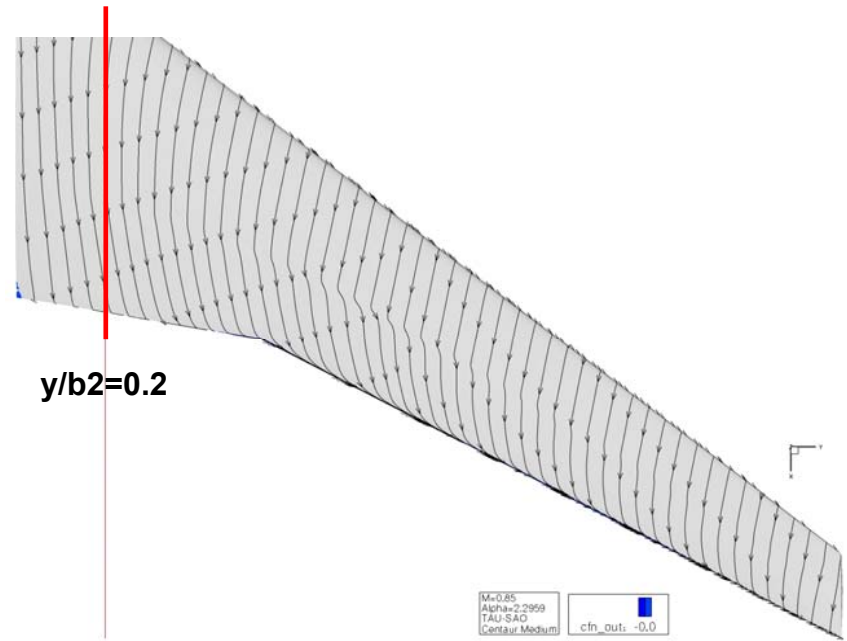
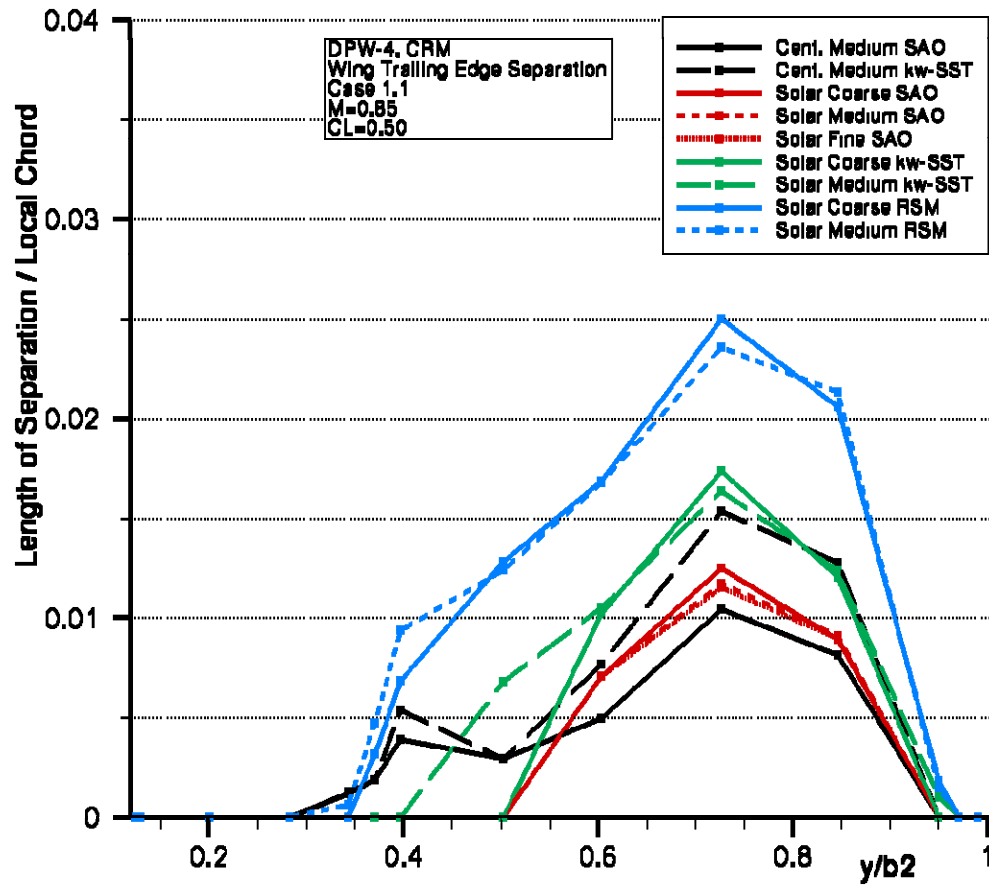


Menter kw-SST, Centaur grid, medium



Case 1.1

— Grid Type/Size, Turbulence Model —





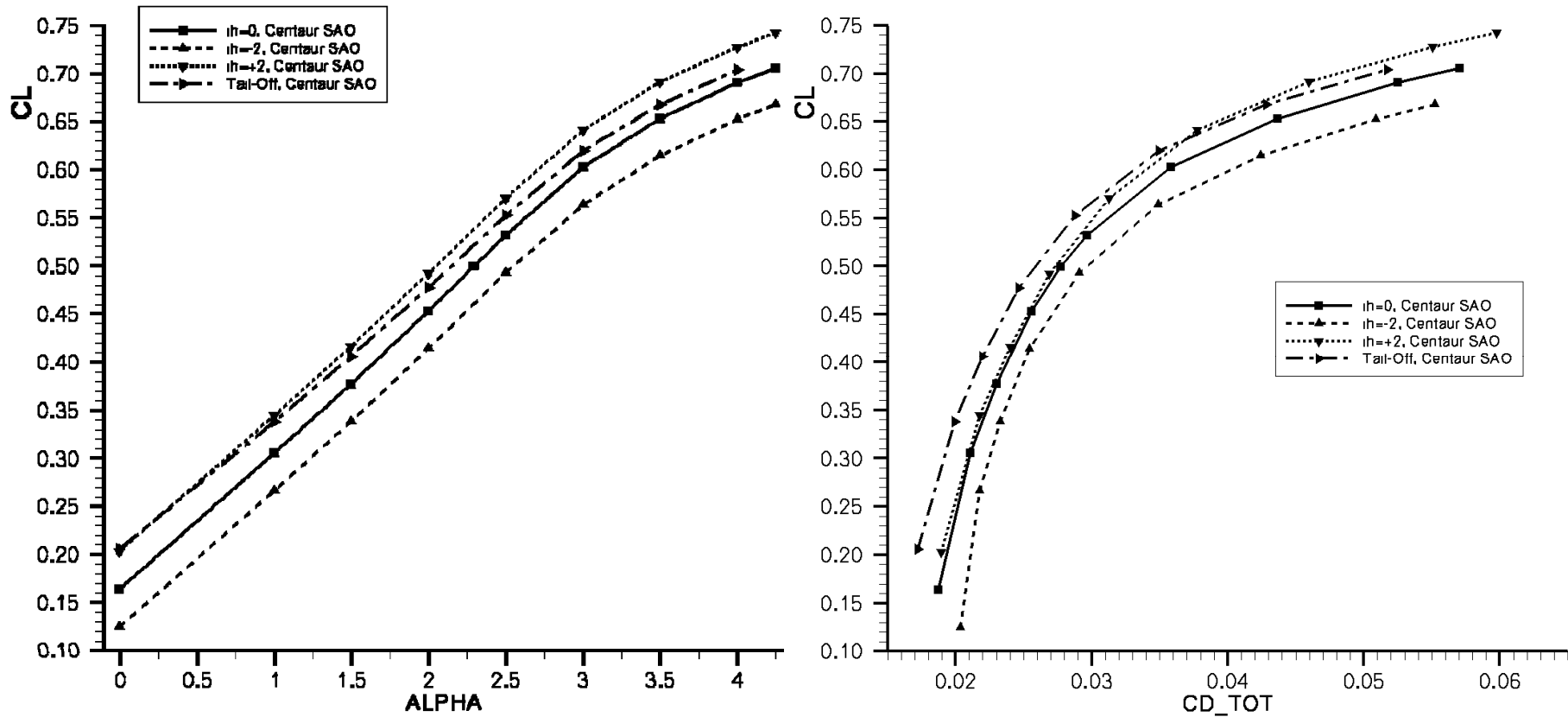
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Case 1.2

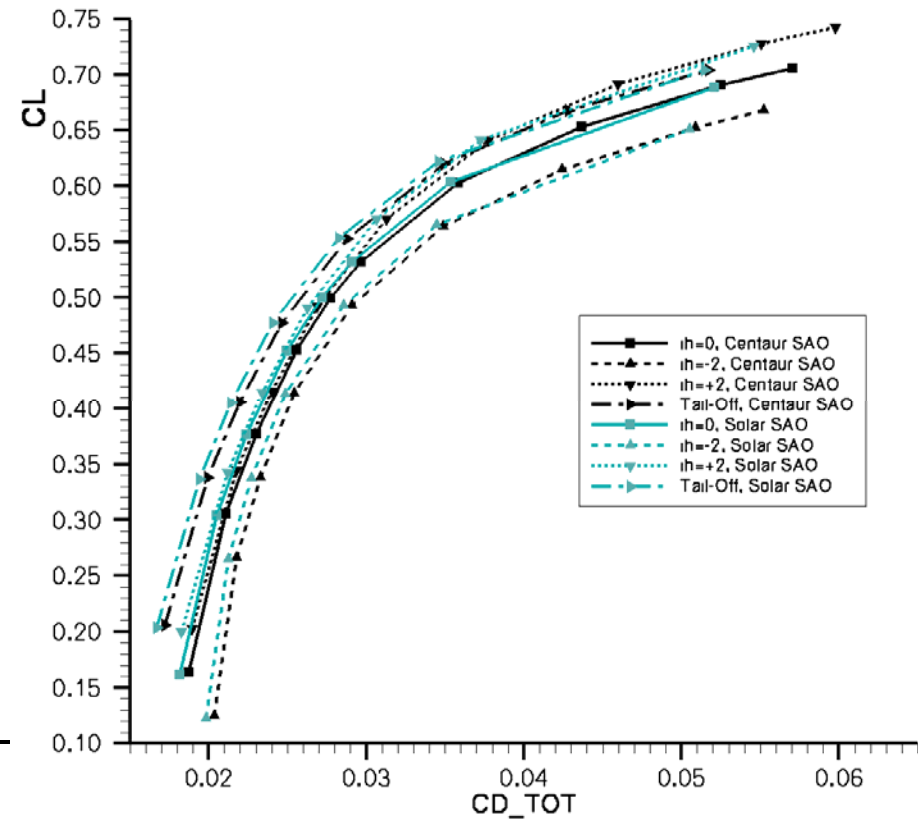
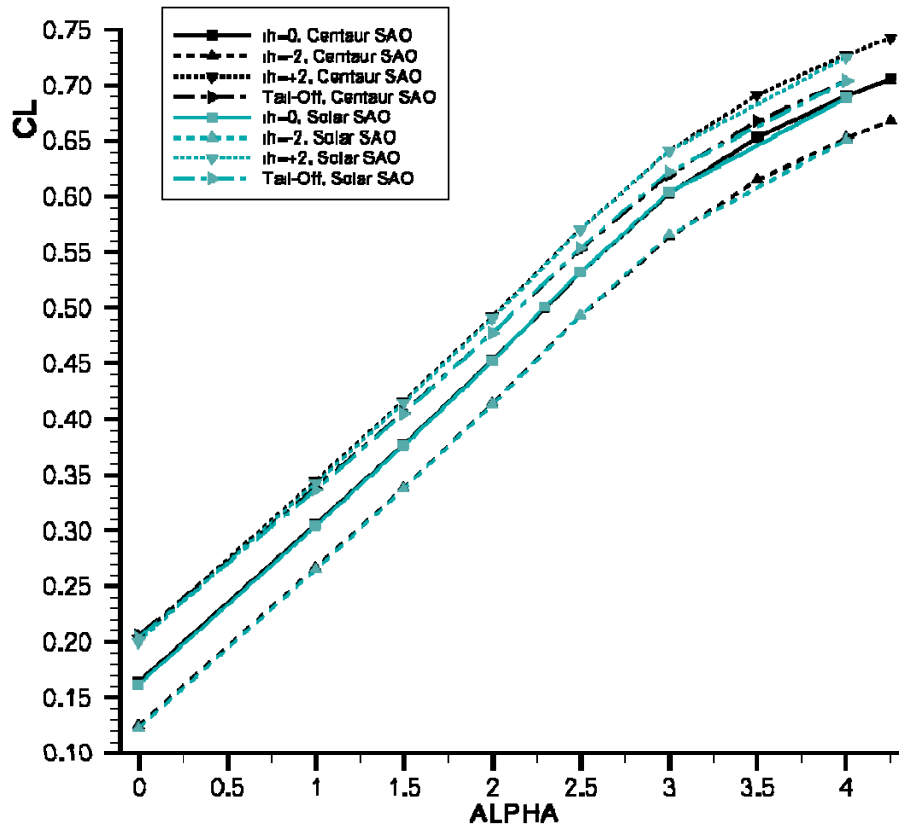
— C_L - α , Polar, HTP Settings —





Case 1.2

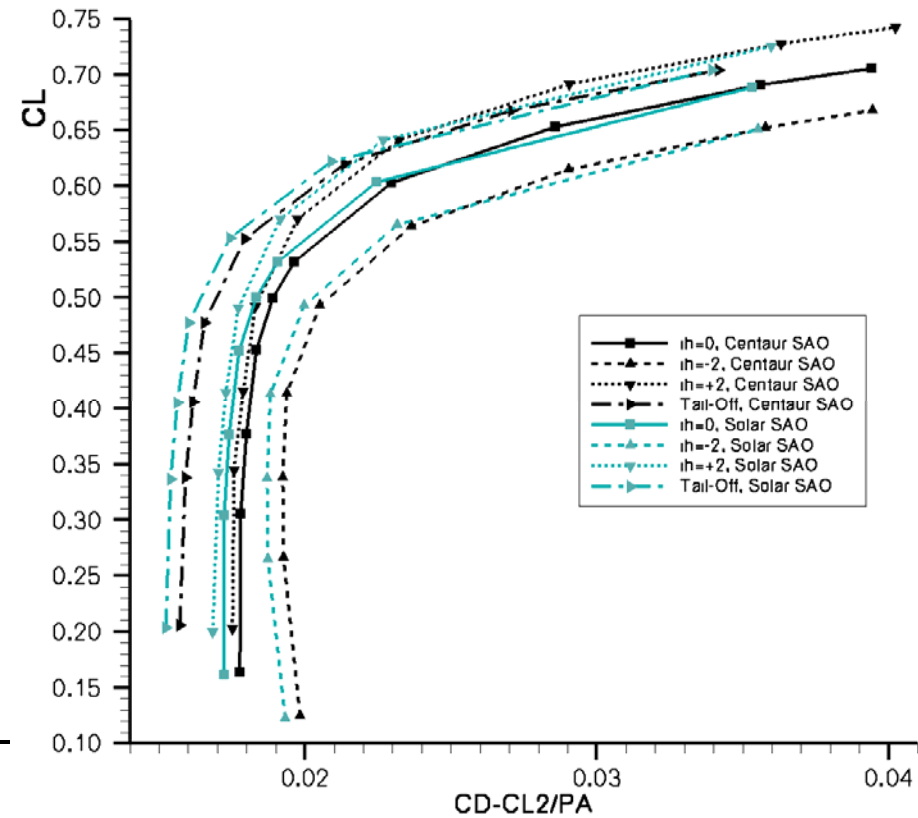
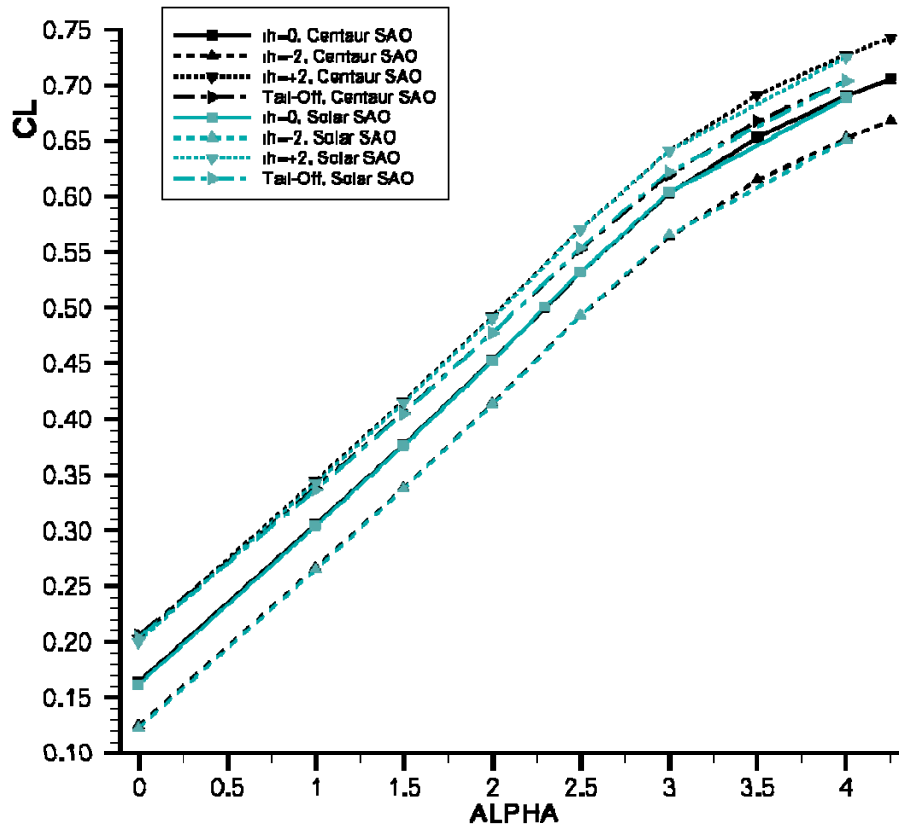
— C_L - α , Polar, HTP Settings, Grid Type —





Case 1.2

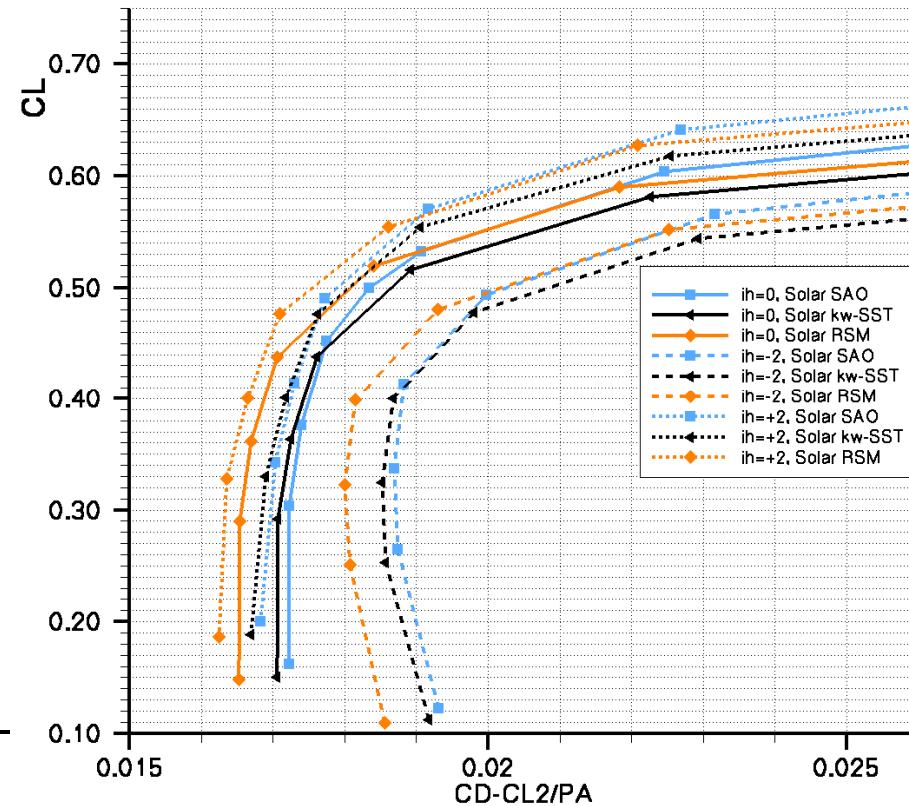
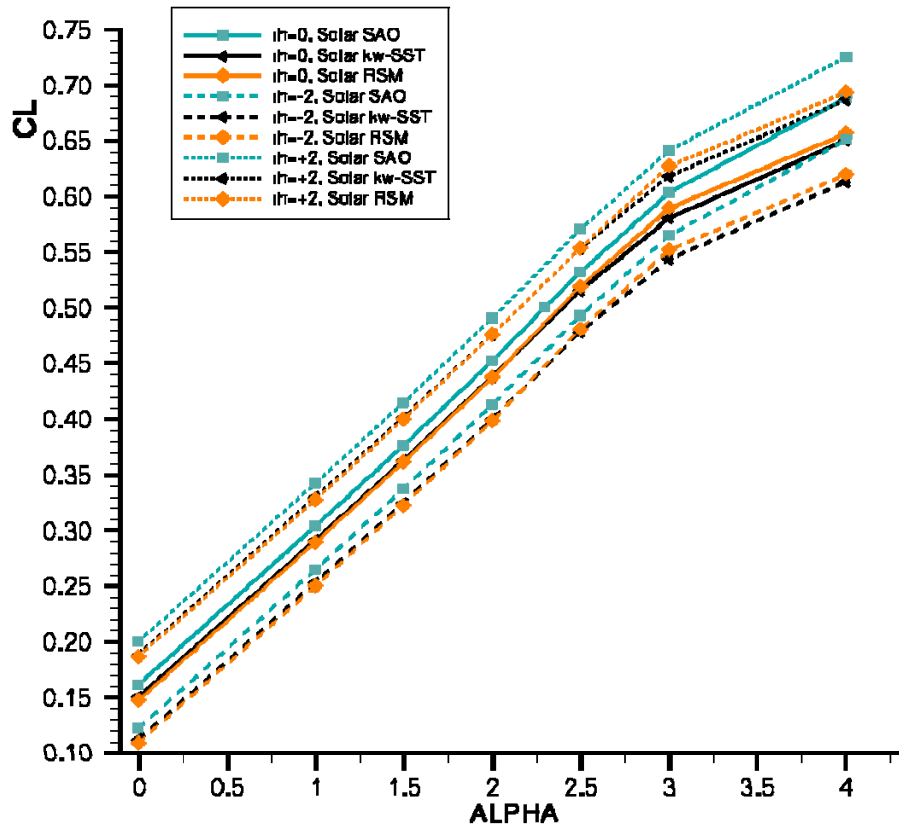
— C_L - α , Polar, HTP Settings, Grid Type —





Case 1.2

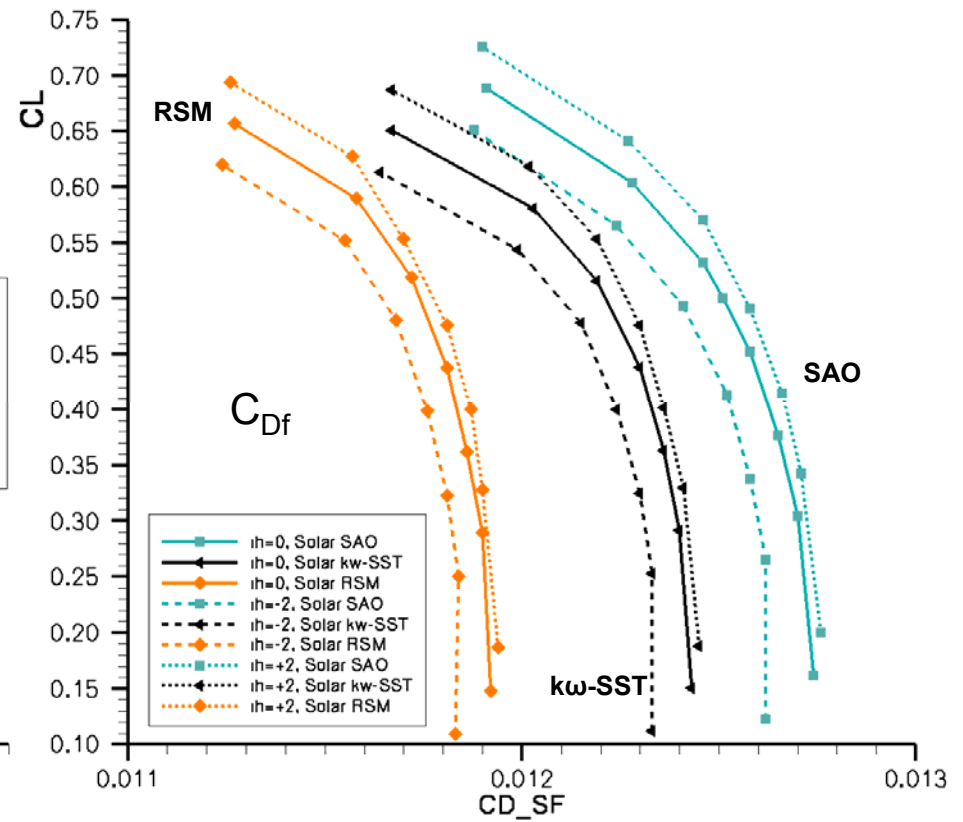
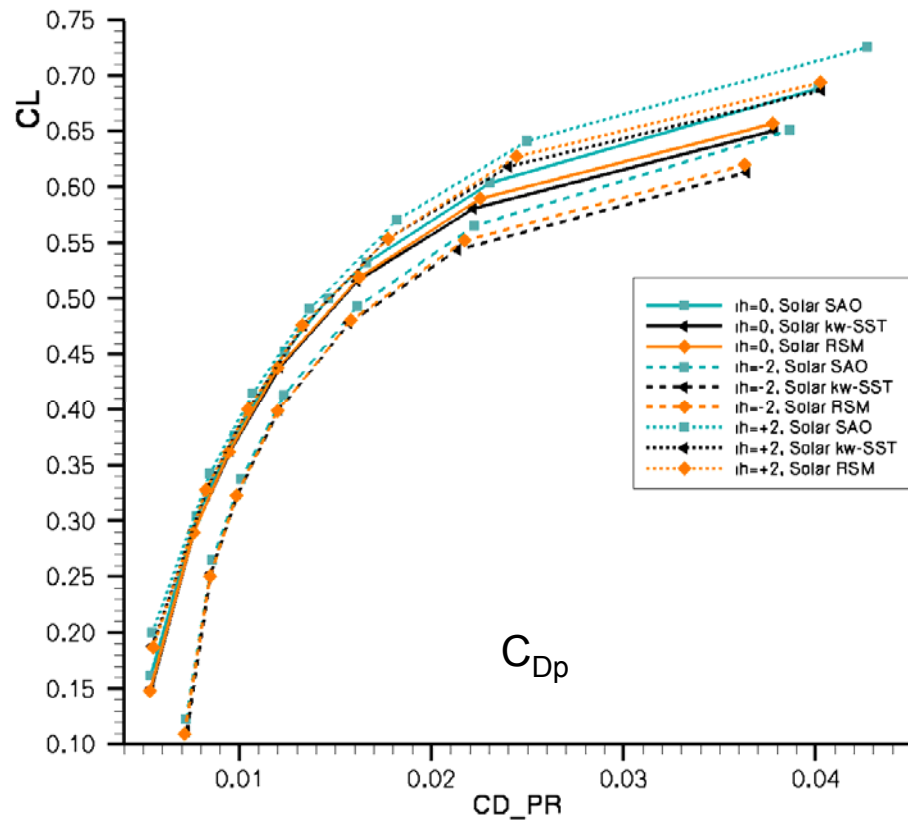
— C_L - α , Polar, Turbulence Model, HTP Setting —





Case 1.2

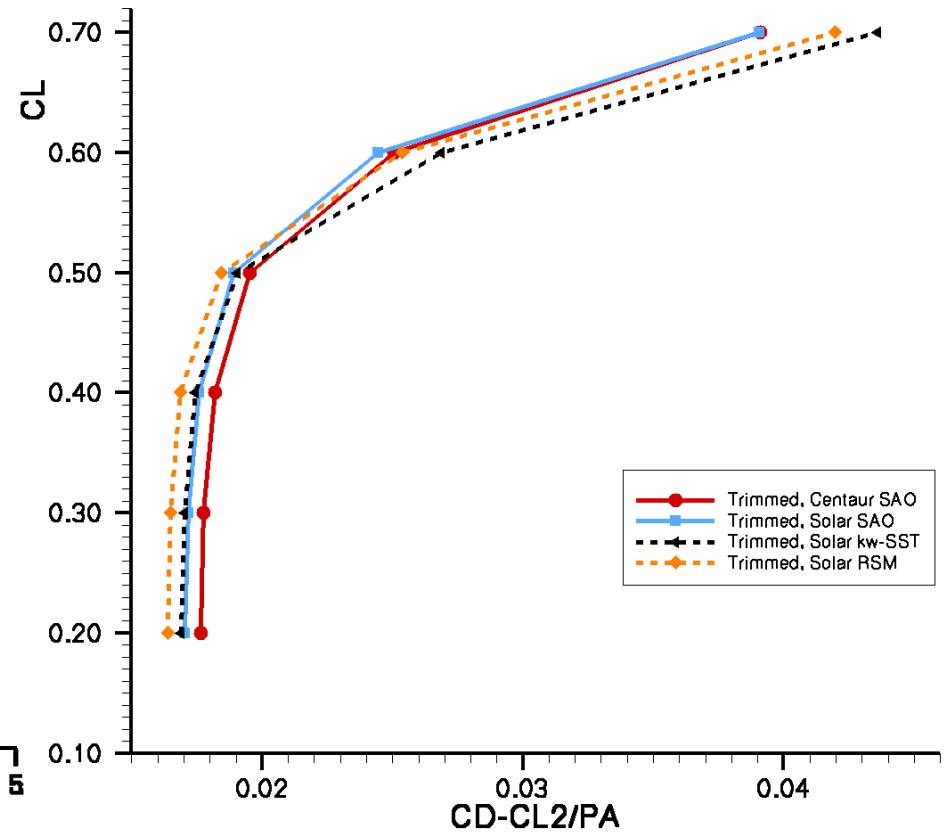
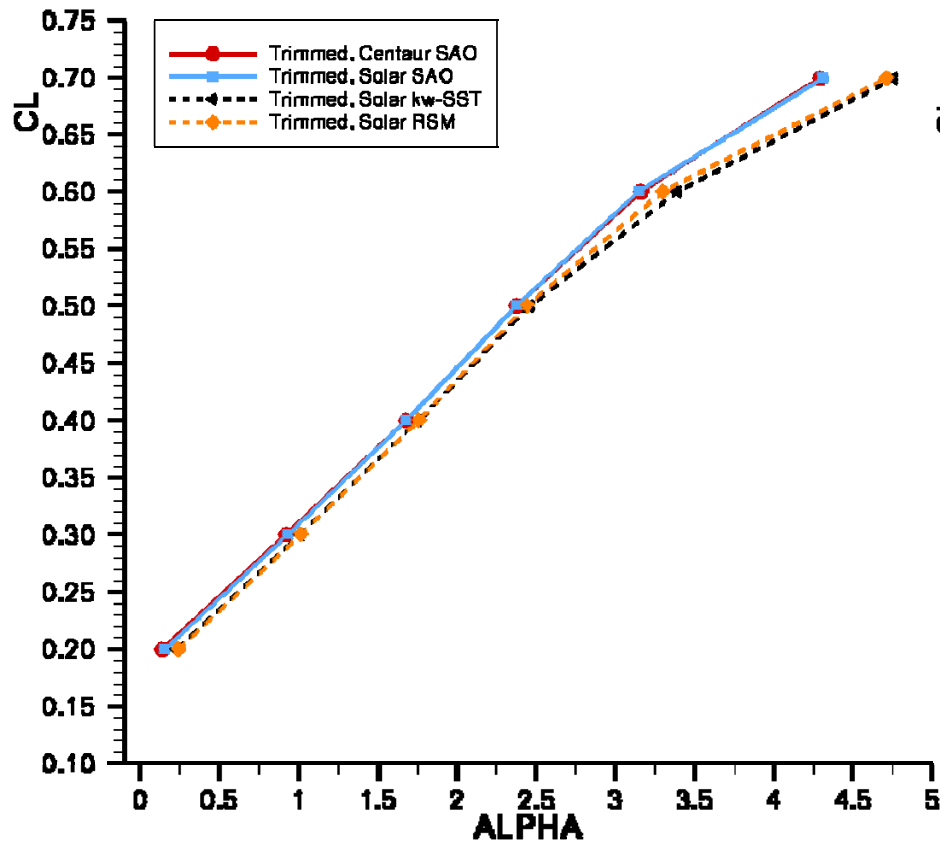
— $C_{Dp/f}$ Polar, Turbulence Model, HTP Setting —





Case 1.2

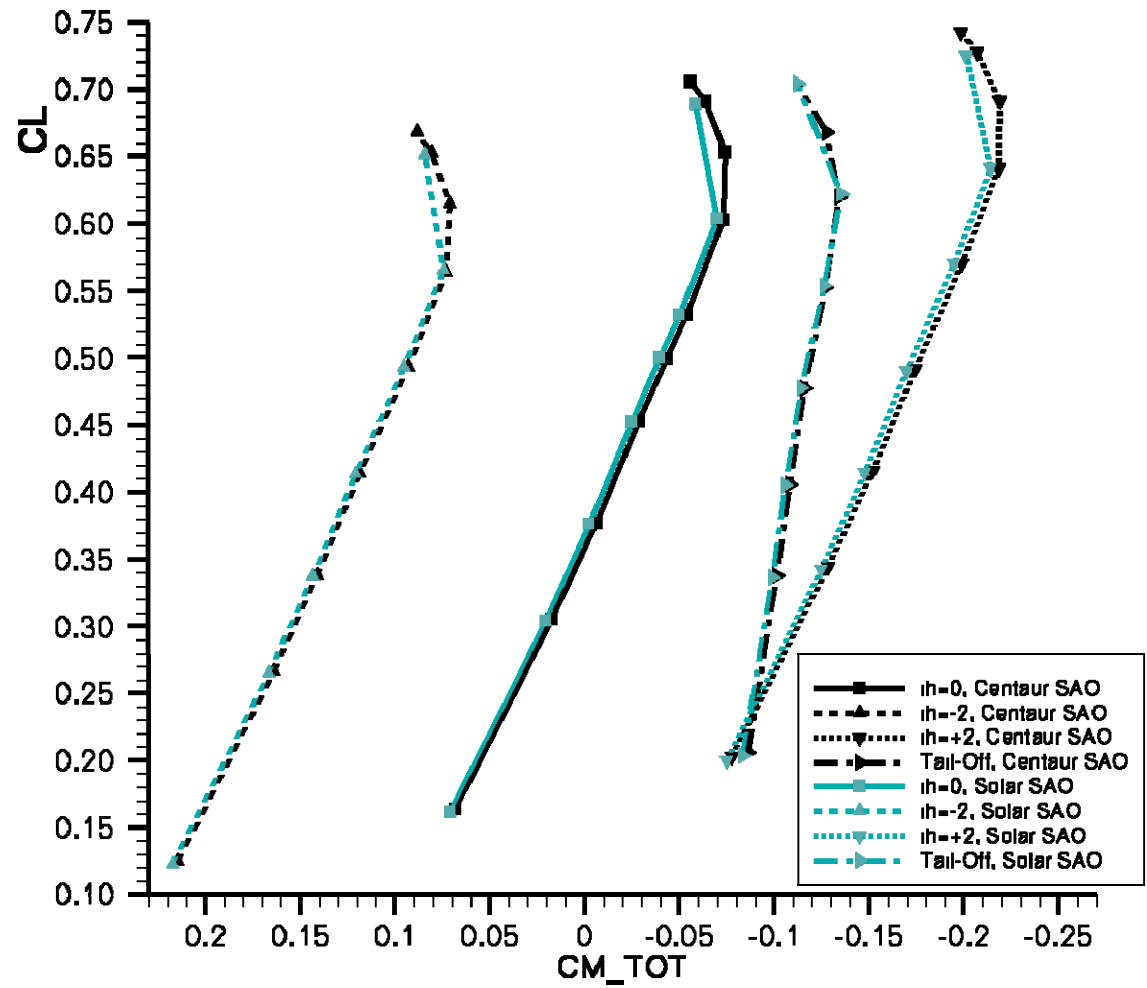
— C_L - α , Polar, Trimmed, Turbulence Model, Grid Type —





Case 1.2

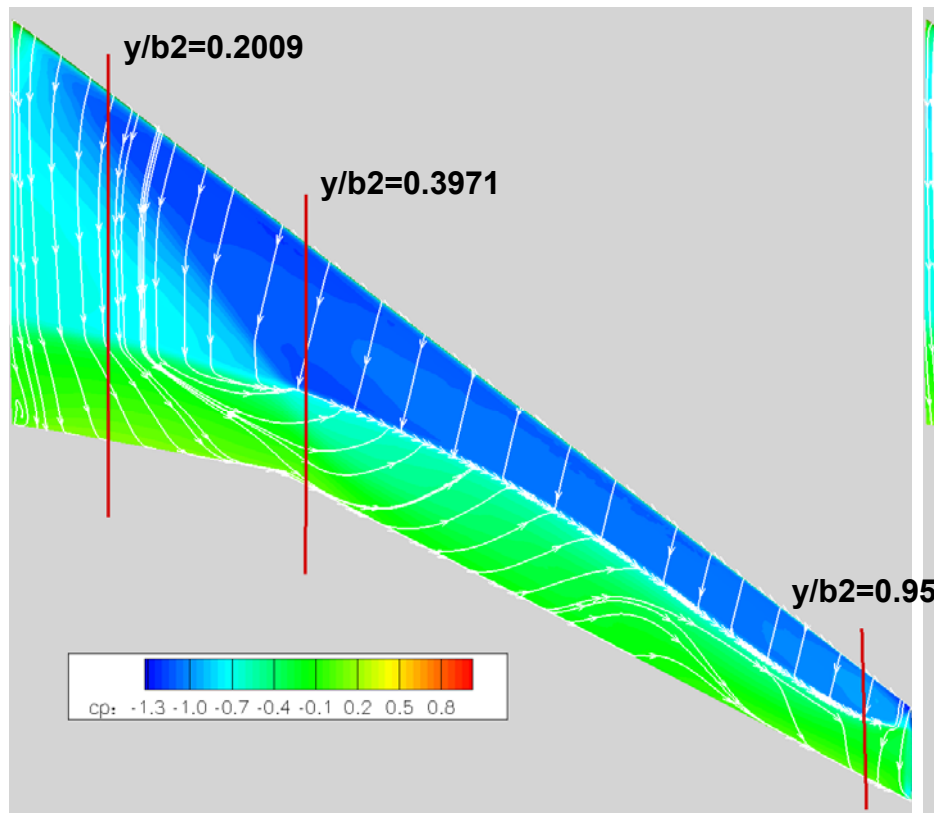
— C_L - C_M , Grid Type —



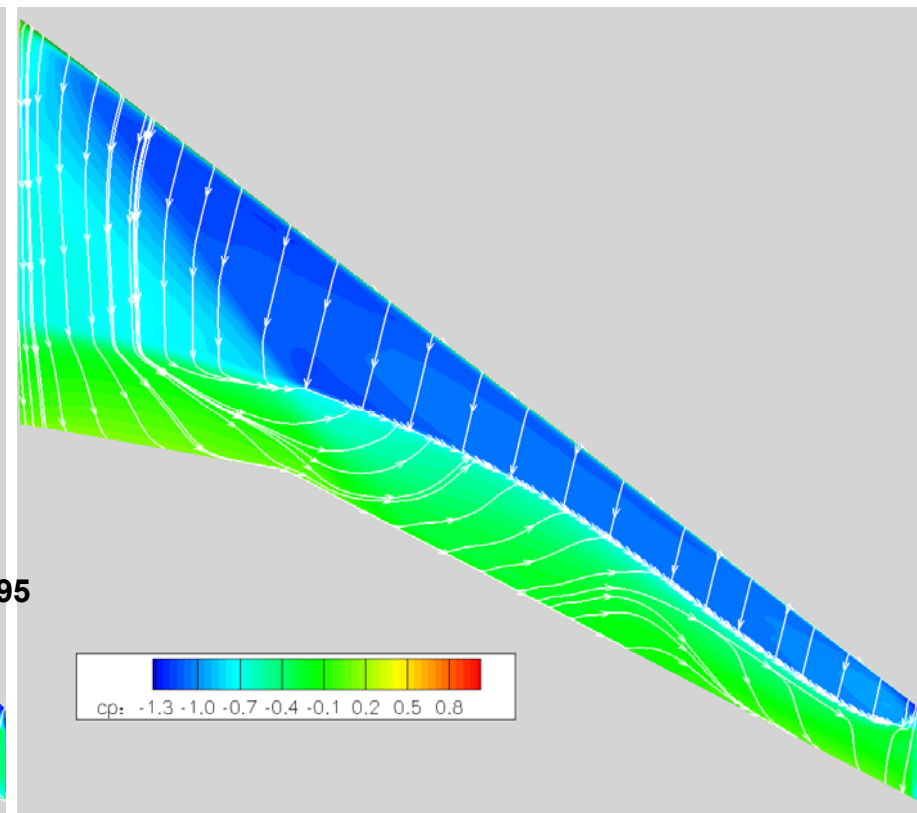


Case 1.2

— Flow Features, $\alpha=4.0^\circ$, Grid Type —



SAO, Centaur medium grid

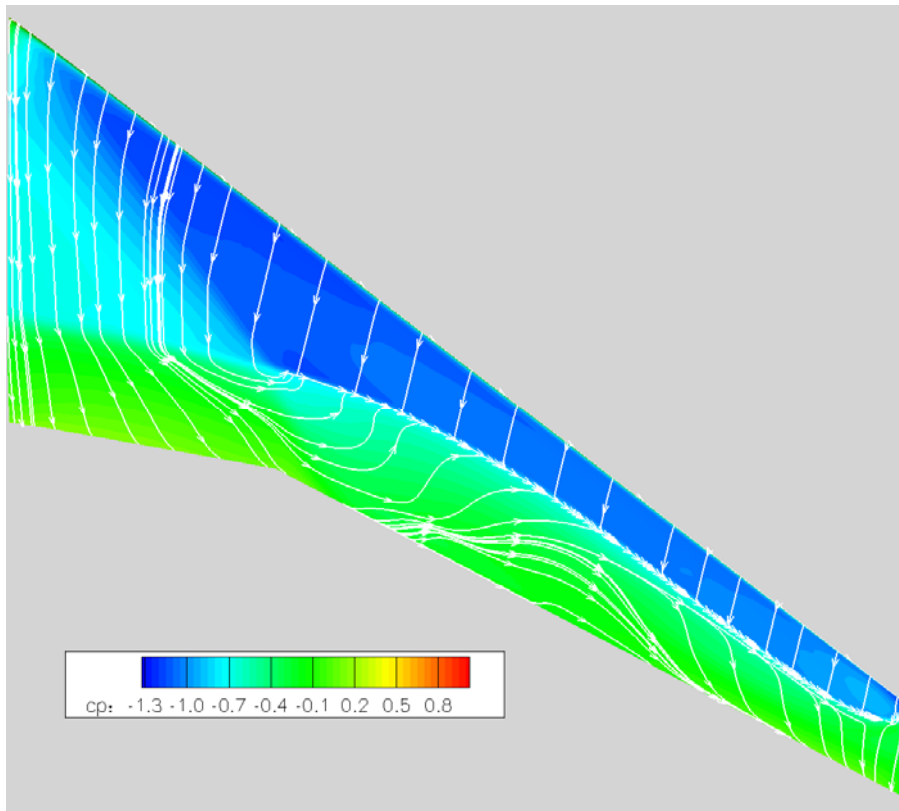


SAO, Solar medium grid

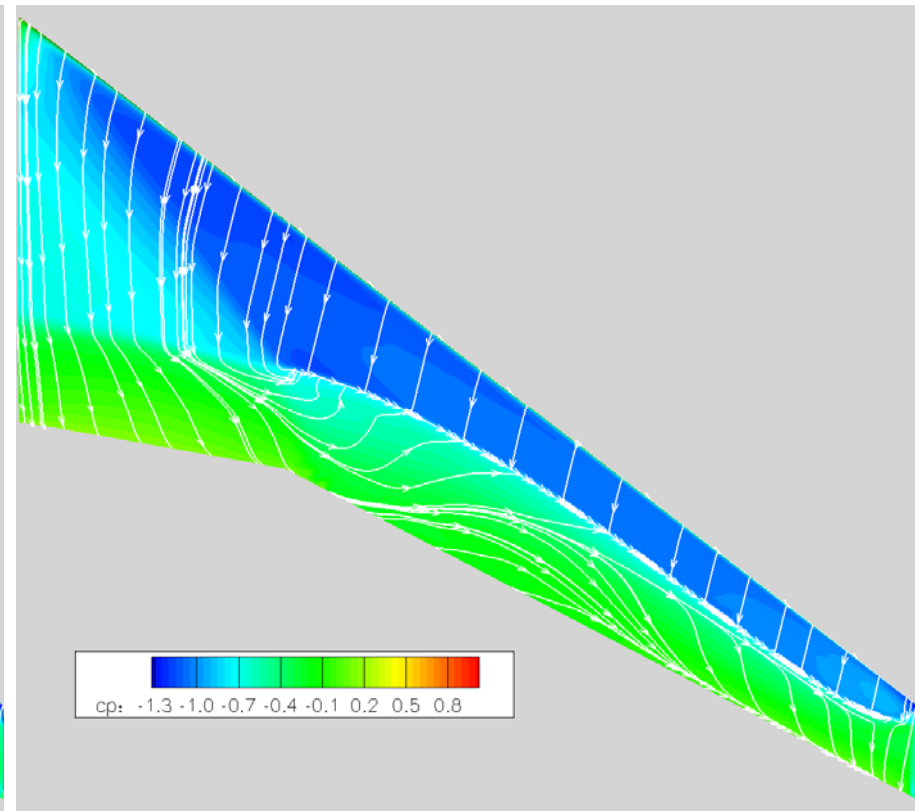


Case 1.2

— Flow Features, $\alpha=4.0^\circ$, Turbulence Model —



Menter $k\omega$ -SST, Solar medium grid

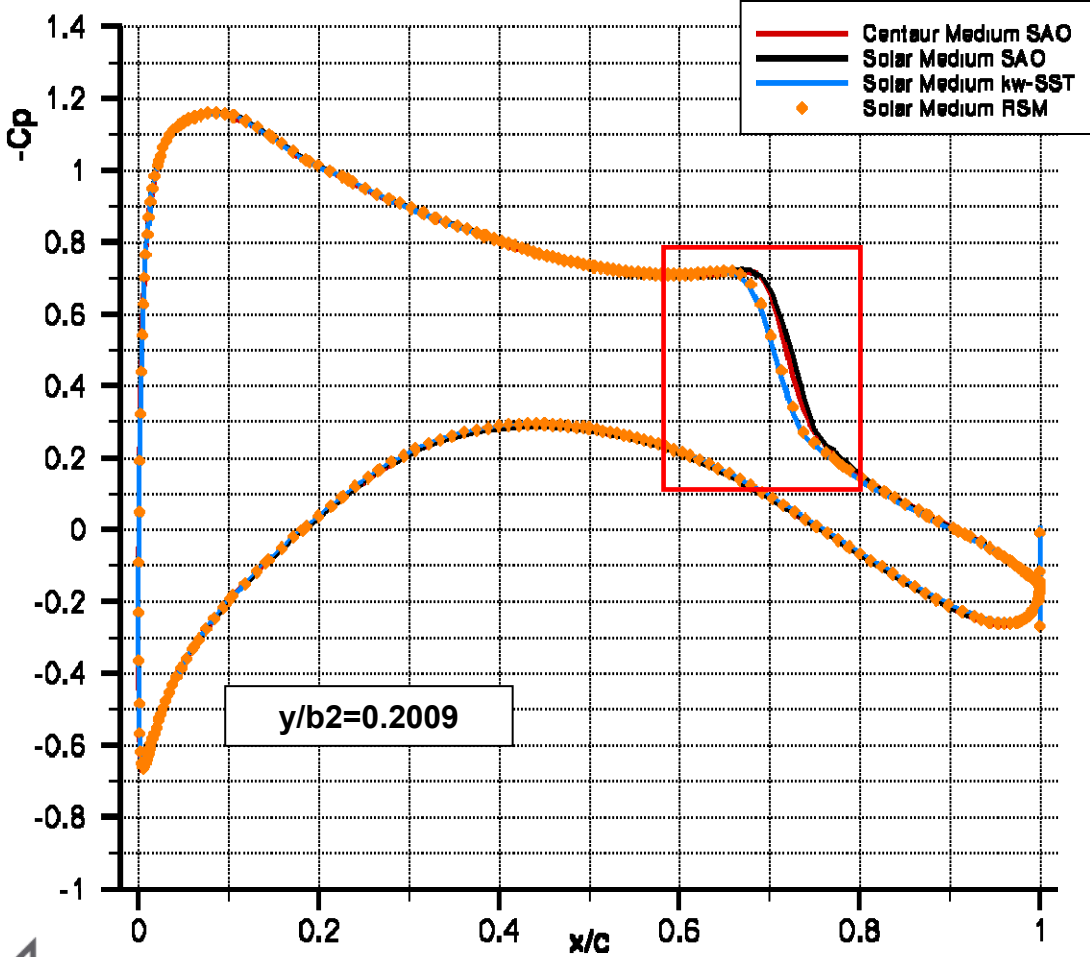


RSM, Solar medium grid



Case 1.2

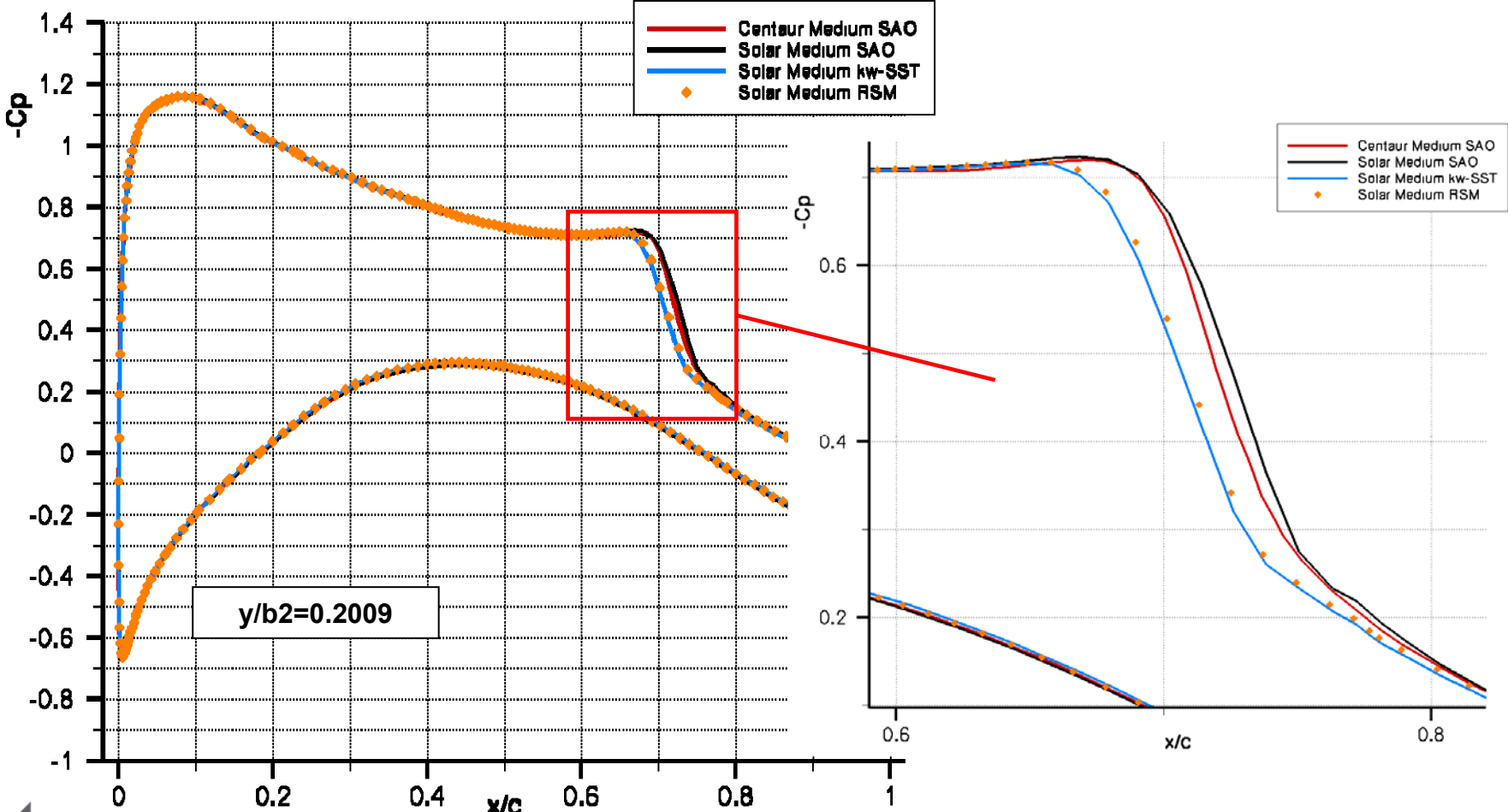
— Pressure Distribution, $\alpha=4.0^\circ$, Turbulence Model —





Case 1.2

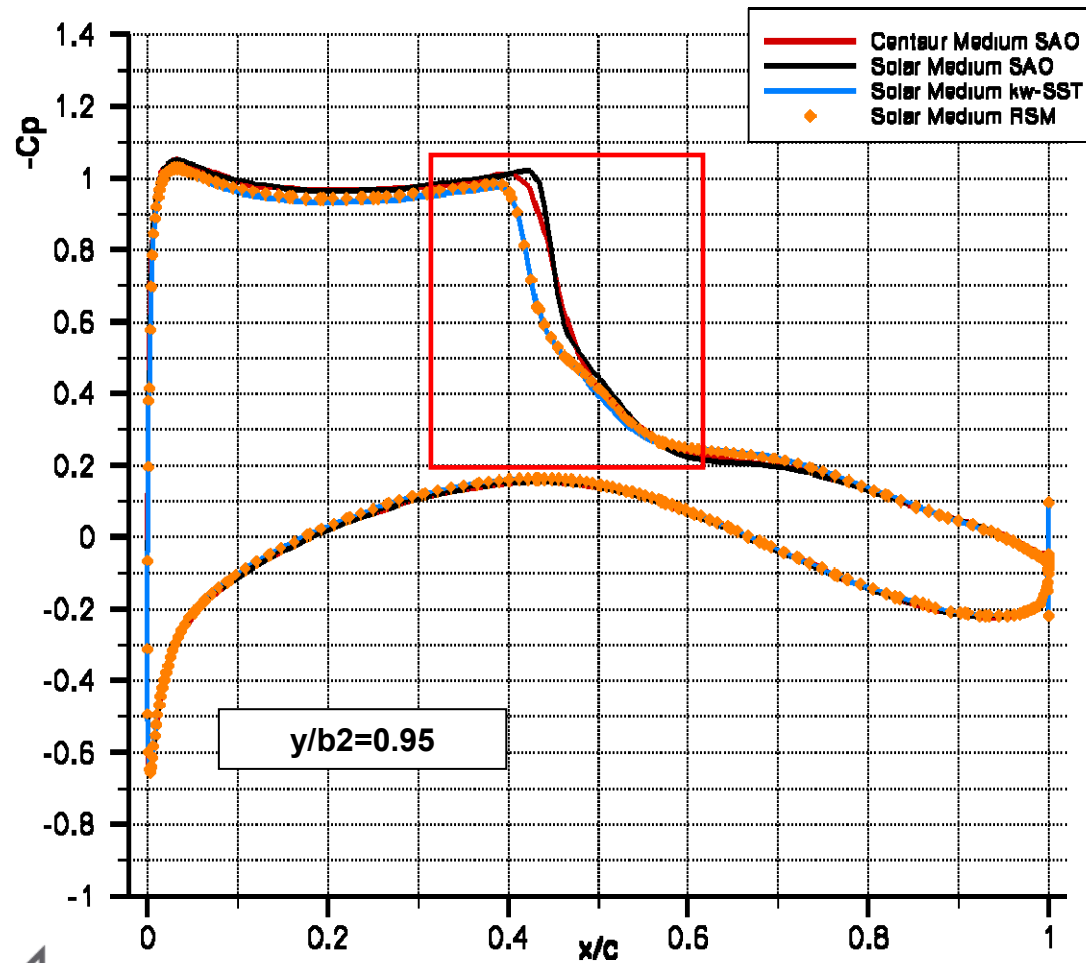
— Pressure Distribution, $\alpha=4.0^\circ$, Turbulence Model —





Case 1.2

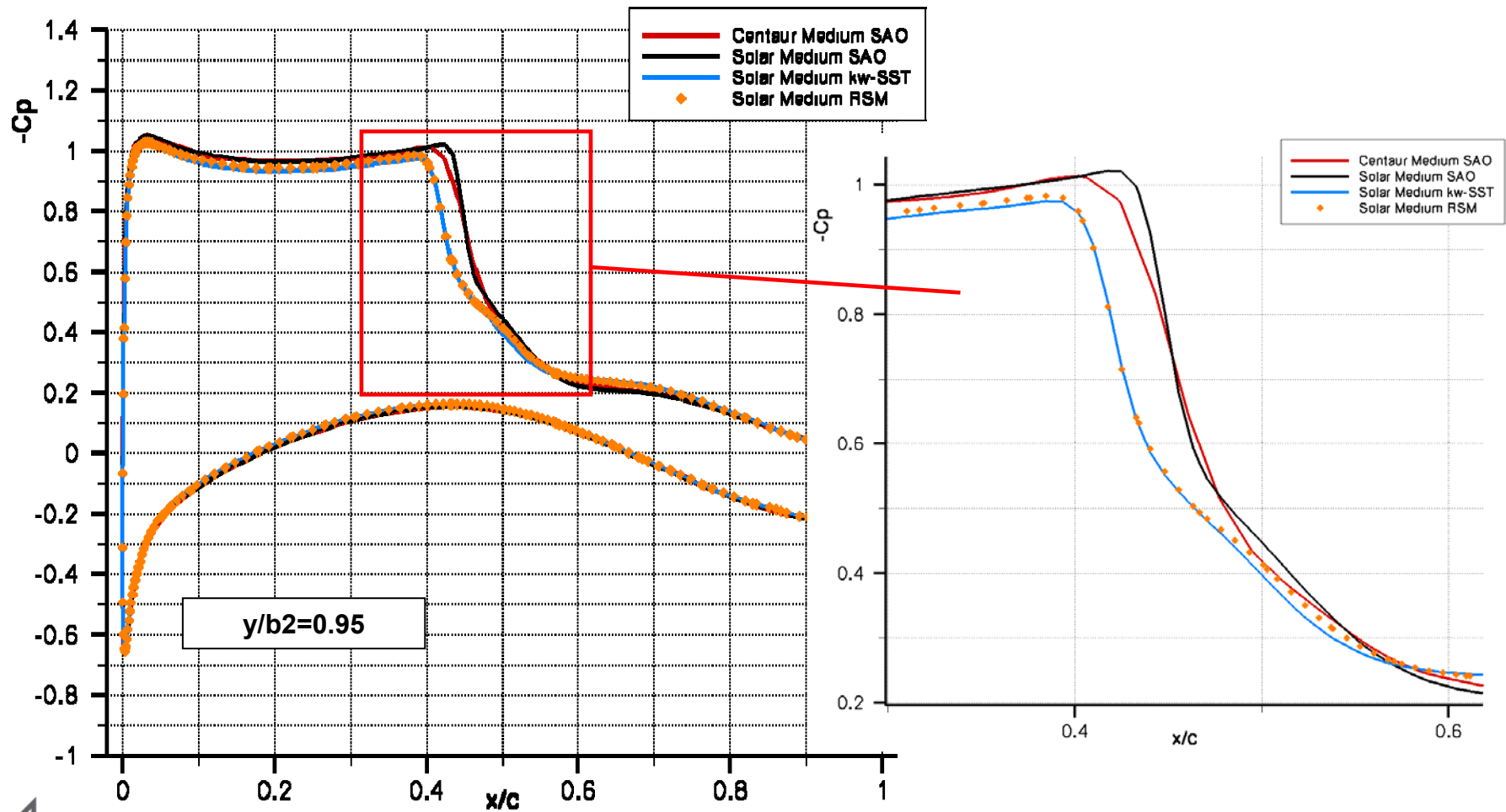
— Pressure Distribution, $\alpha=4.0^\circ$, Turbulence Model —





Case 1.2

— Pressure Distribution, $\alpha=4.0^\circ$, Turbulence Model —

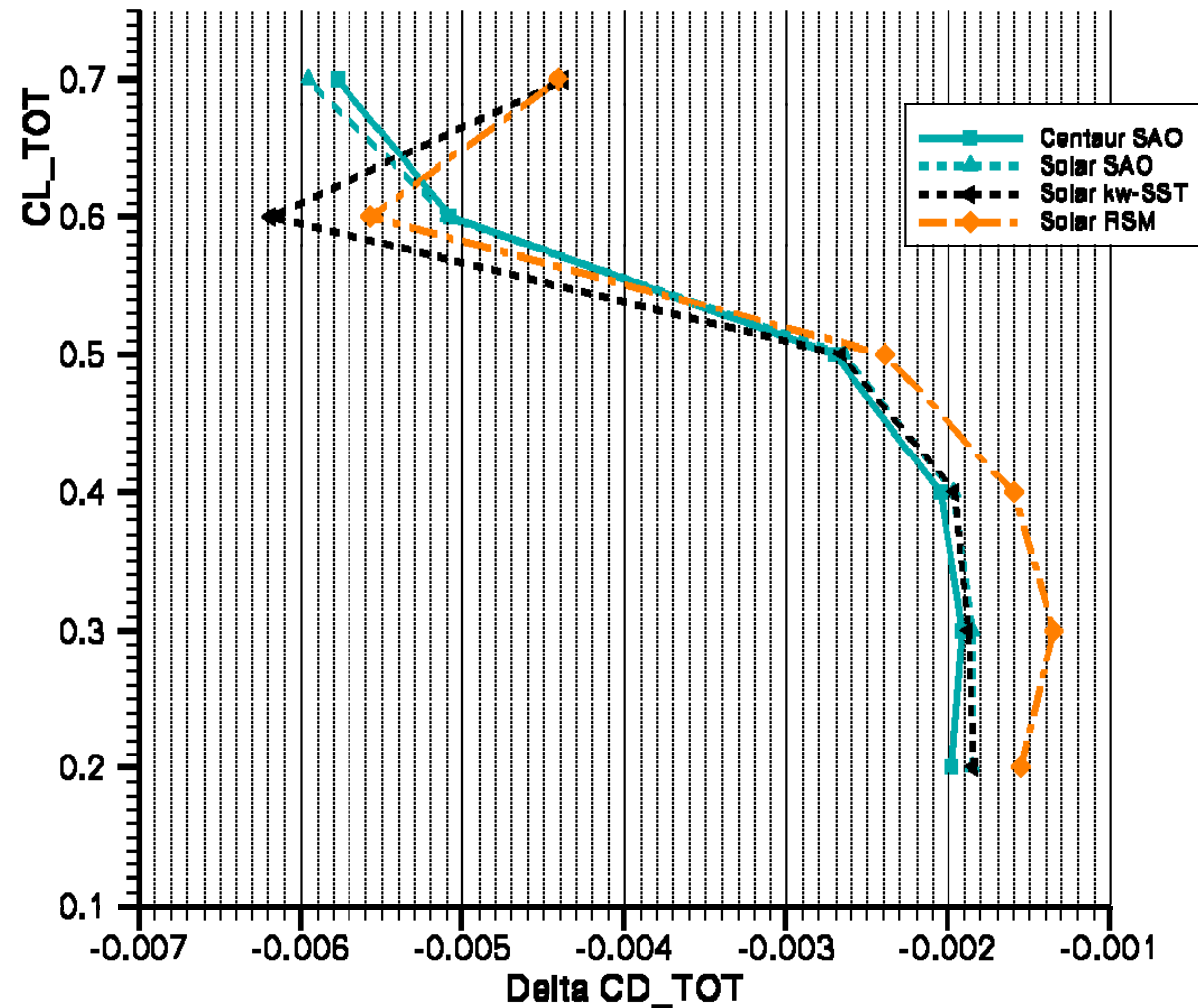




Case 1.2

— Delta Drag, Turbulence Model —

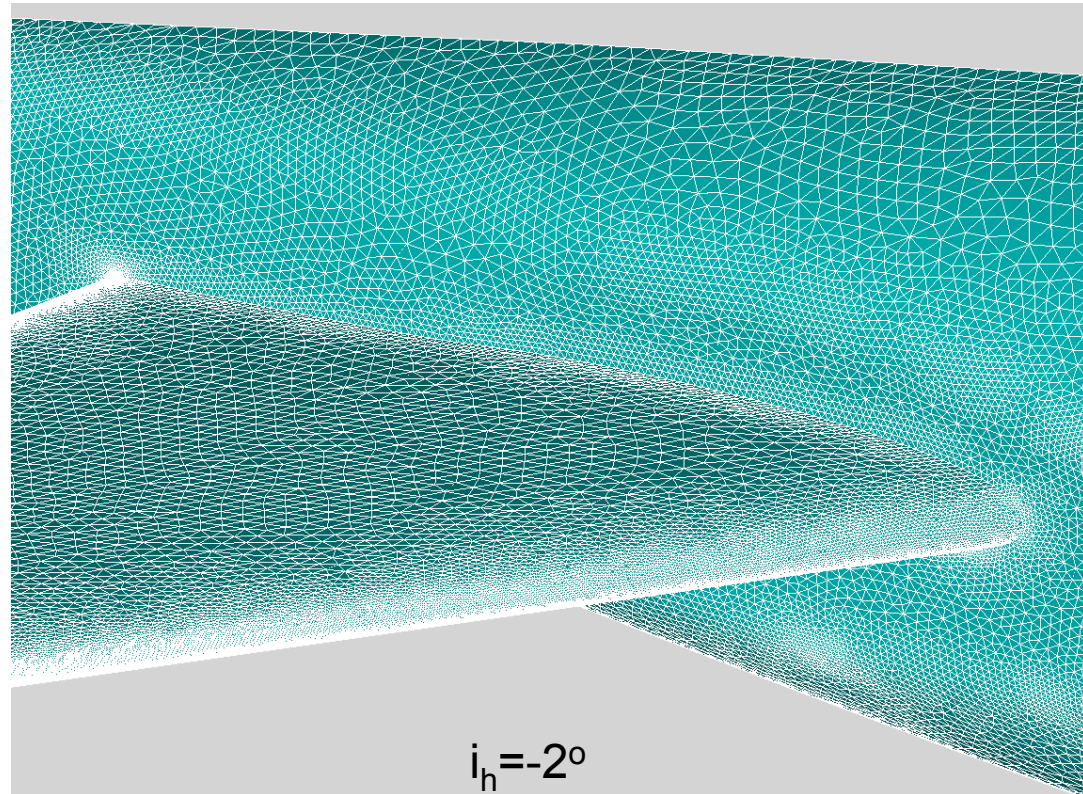
➤ Tail-off minus trimmed configuration



Case 1.2

— HTP Setting Modification —

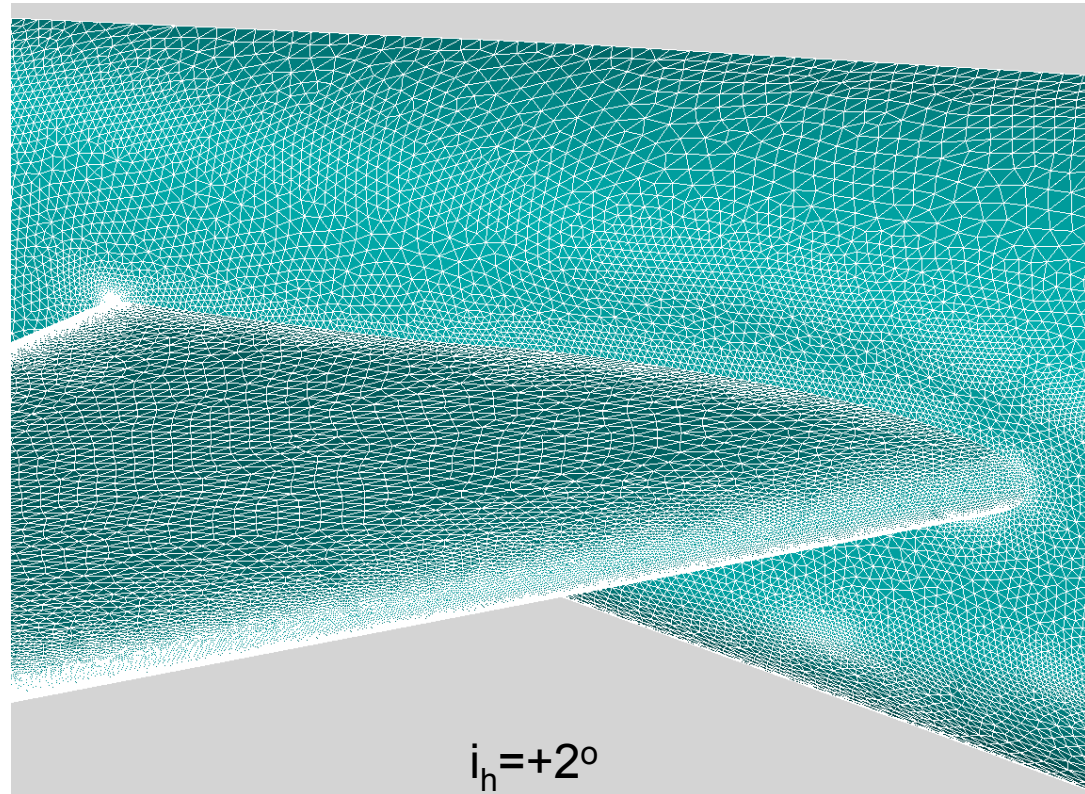
- Iterative modification of i_h during CFD calculations towards $C_M=0$, $C_L=0.5$
- Mesh deformation used based on radial basis functions
- Differences to interpolated data are small
- $\Delta i_h = 0.0029^\circ$
- $\Delta \alpha = 0.000018^\circ$



Case 1.2

— HTP Setting Modification —

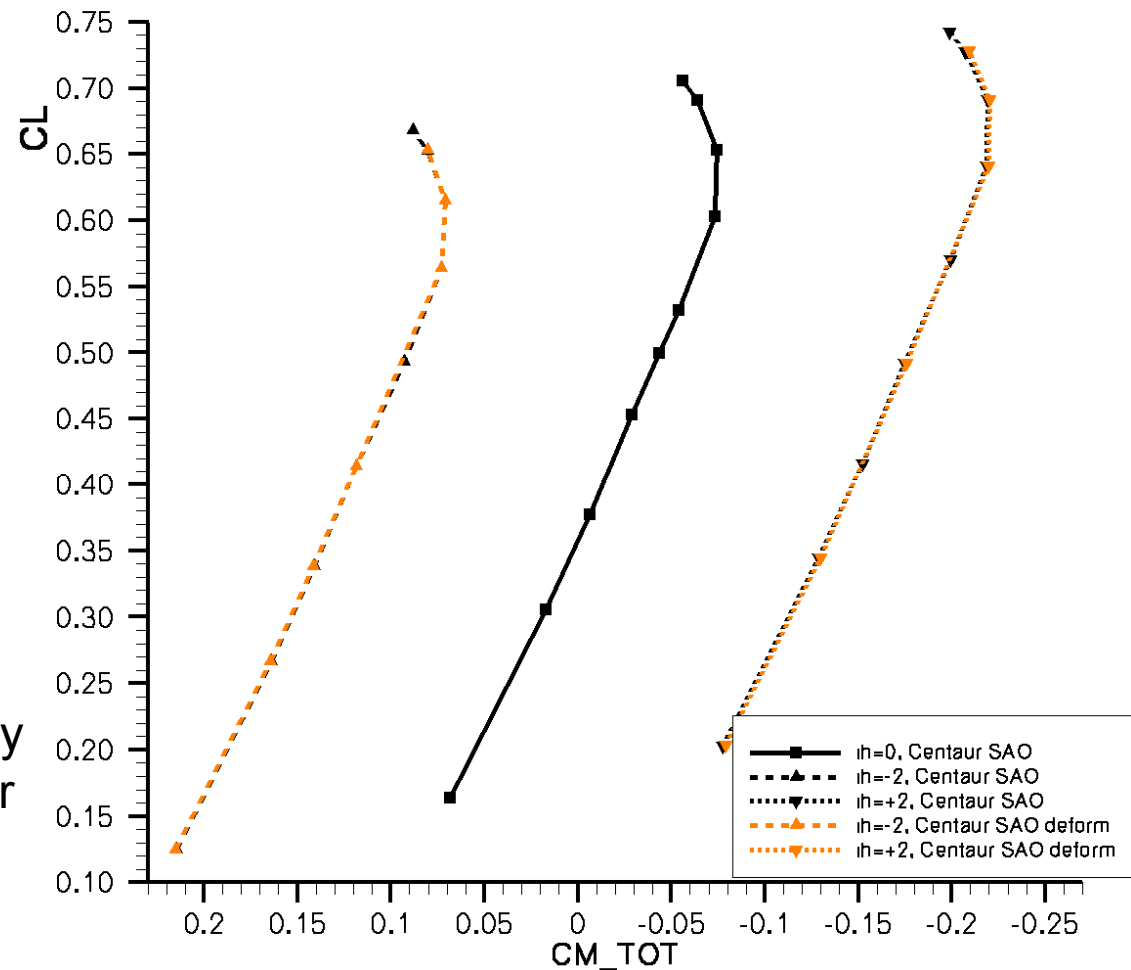
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Case 1.2

— HTP Setting Modification —

- Iterative modification of i_h during CFD calculations towards $C_M=0$, $C_L=0.5$
- Mesh deformation used based on radial basis functions
- Differences to interpolated data are small
- $\Delta i_h=0.0029^\circ$
- $\Delta\alpha=0.000018^\circ$
- C_M for deformed grids nearly identical to C_M calculated for separately generated grids.





Content

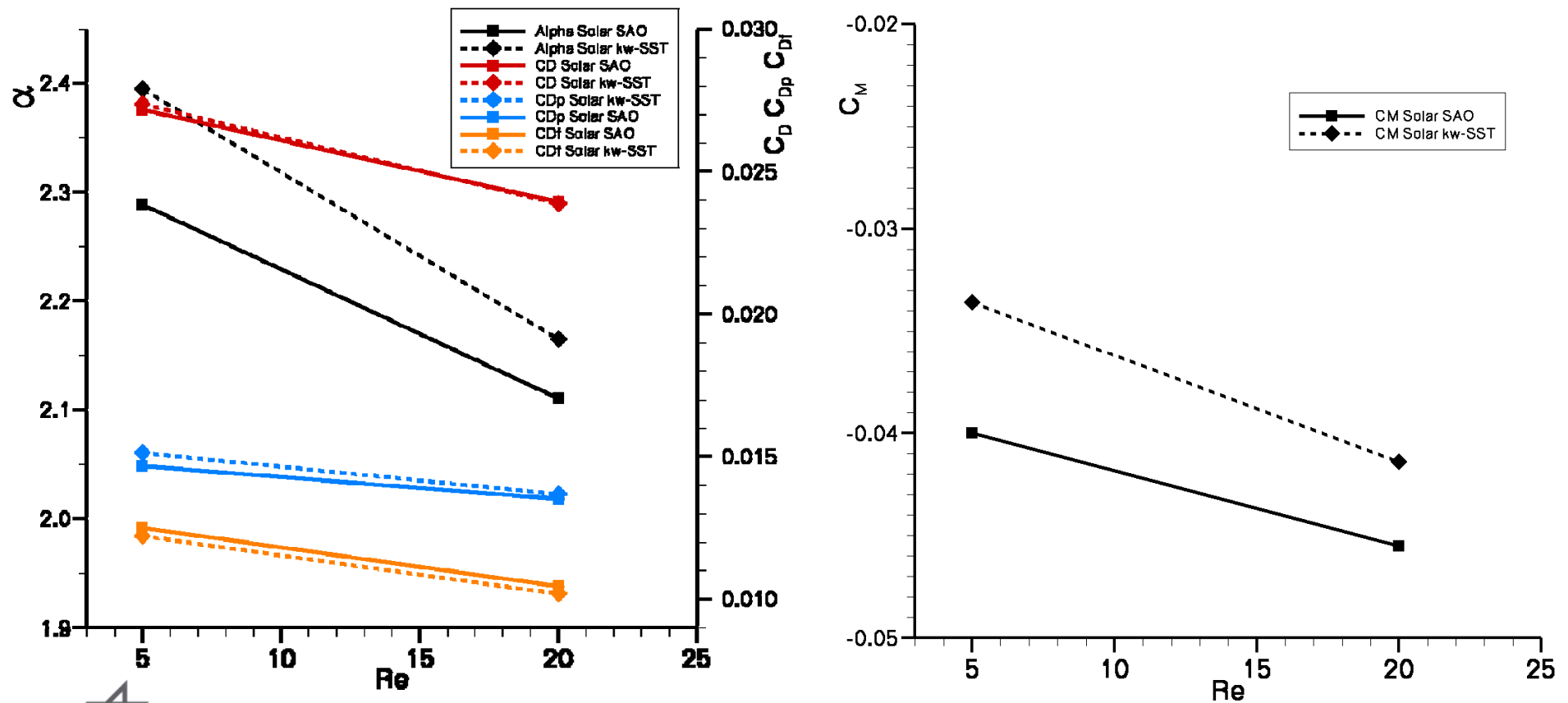
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Case 3

— Re Influence, Turbulence Model —

$\Delta\alpha \approx 8-10\%$, $\Delta C_D \approx 12-13\%$, $\Delta C_M \approx 14-23\%$





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Conclusions (Preliminary)

— Lessons Learned —

- Solar quad-dominant unstructured surface meshes and hex-BL-resolution technique produce very good hierarchy of grids with high quality.
- Better leading edge and shock resolution with less overall nodes achievable.
- Aerodynamic coefficients and deltas for medium Centaur/Solar grids are similar.
- Grid refinement indicates a nearly linear behaviour for $1/N^{2/3}$.
- Wing fuselage separation is influenced by the grid type/size.
- Trailing edge separation size is mainly influenced by the turbulence model.
- Trimmed polars: the grid influence is less important than the turbulence model.
- Iterative setting variation of HTP in CFD loop towards $C_M=0$ for $C_L=0.5$:
 - Very small differences of i_h and α for trimmed configuration compared to interpolation method based on results from separately generated grids.